

The weak interaction

Part 3 & 4

- The K^0 - \bar{K}^0 system
- The CKM mechanism and the unitarity triangle
- Neutrinos

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Achille Stocchi
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The $K^0-\bar{K}^0$ system

Remember the strange particles ?

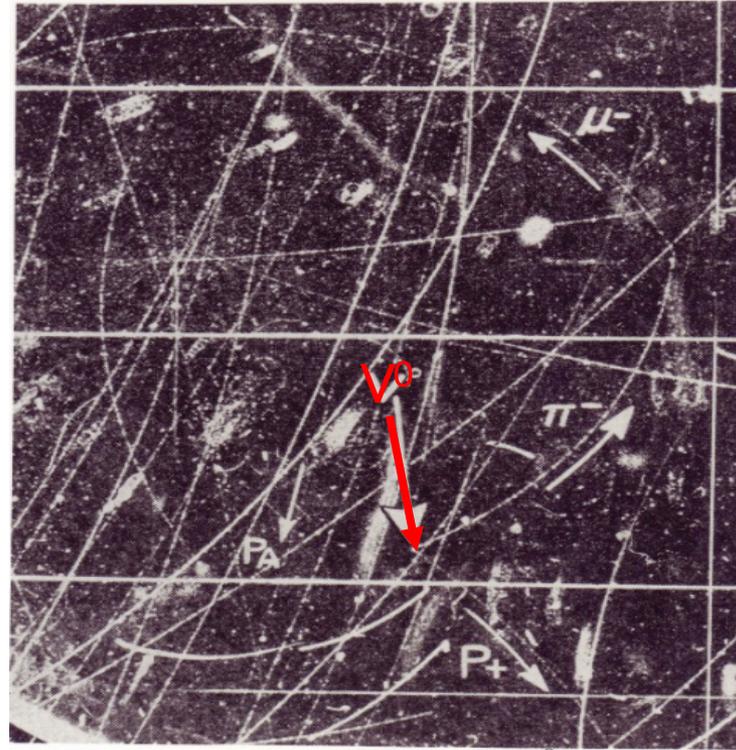
Observation of Long-Lived Neutral V Particles*

K. LANDE, E. T. BOOTH, J. IMPEDUGLIA, AND L. M. LEDERMAN,
Columbia University, New York, New York

AND

W. CHINOWSKY, Brookhaven National Laboratory,
Upton, New York

(Received July 30, 1956)



2

Brookhaven, 1956

→ Two states :

$$|K_1\rangle \rightarrow \pi\pi$$

$$|K_2\rangle \rightarrow \pi\pi\pi$$

Same mass (~ 500 MeV)

Very different lifetimes

K_2 lifetime ~ 10000 K_1 lifetime due to phase space

$$M(\pi) \sim 140 \text{ MeV}$$

$$M(K) \sim 500 \text{ MeV}$$

CP violation in the K^0 system

$$|K^0\rangle = |\bar{s}d\rangle$$

$$|\bar{K}^0\rangle = |\bar{d}s\rangle$$

$$\text{CP} |K^0\rangle = |\bar{K}^0\rangle \Rightarrow$$

$$|K^0\rangle = |\bar{s}d\rangle$$

$$|\bar{K}^0\rangle = |\bar{d}s\rangle$$

not CP eigenstates

One can build :

$$|K_1\rangle = \frac{1}{\sqrt{2}} (|K^0\rangle + |\bar{K}^0\rangle)$$

$$|K_2\rangle = \frac{1}{\sqrt{2}} (|K^0\rangle - |\bar{K}^0\rangle)$$

CP eigenstates

$$|K_1\rangle \rightarrow \pi\pi$$

$$|K_2\rangle \rightarrow \pi\pi\pi$$

$$\text{CP}(\pi\pi) = +1 \text{ and } \text{CP}(\pi\pi\pi) = -1$$

if CP is a good symmetry for the weak interaction : ~~$|K_2\rangle \rightarrow \pi\pi$~~

$$|K_1\rangle \rightarrow \pi\pi$$

After some time, pure K_2 beam

initial beam
 K_1 and K_2

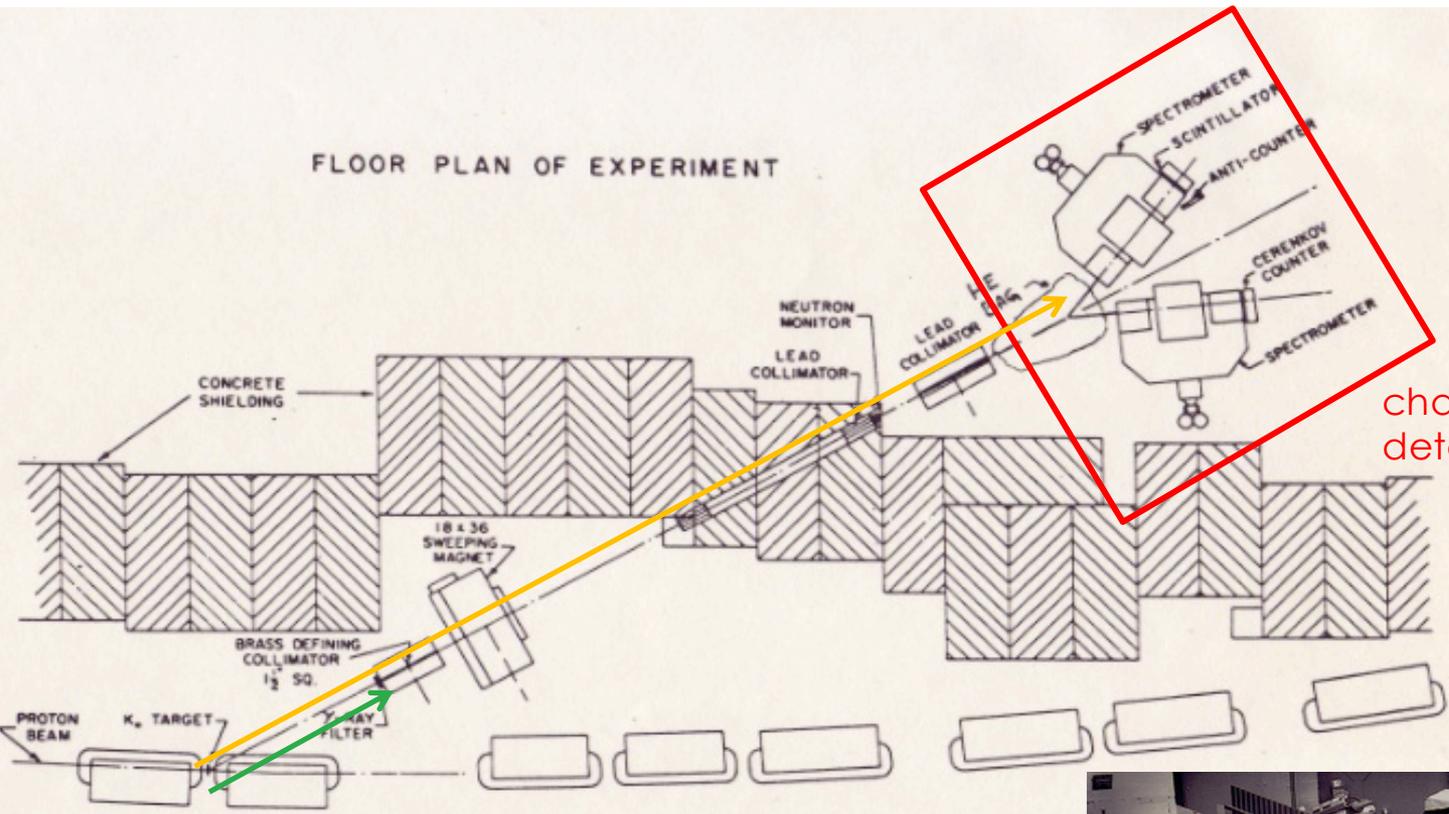


Search for the signal of the decay $|K_2\rangle \rightarrow \pi\pi$ far (20 meters)
from the production point of the K_1 and K_2

?

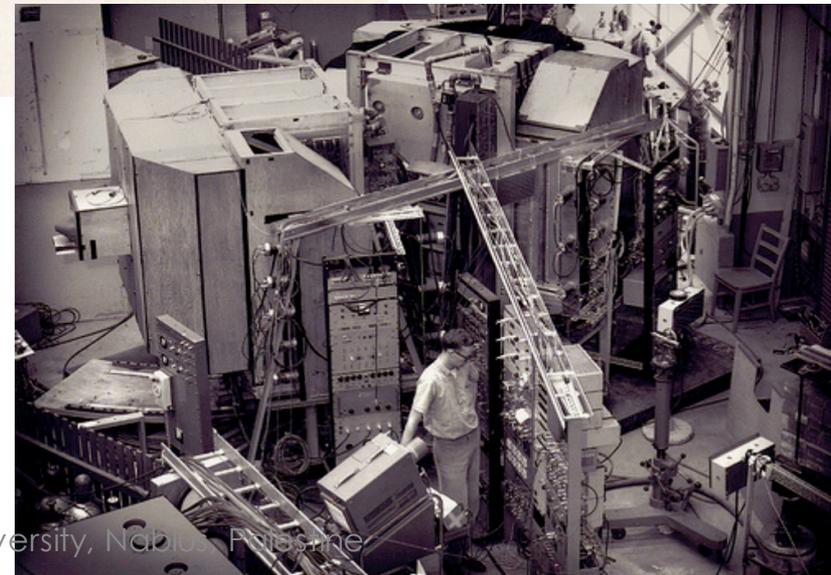
Cronin & Fitch experiment 1964

FLOOR PLAN OF EXPERIMENT

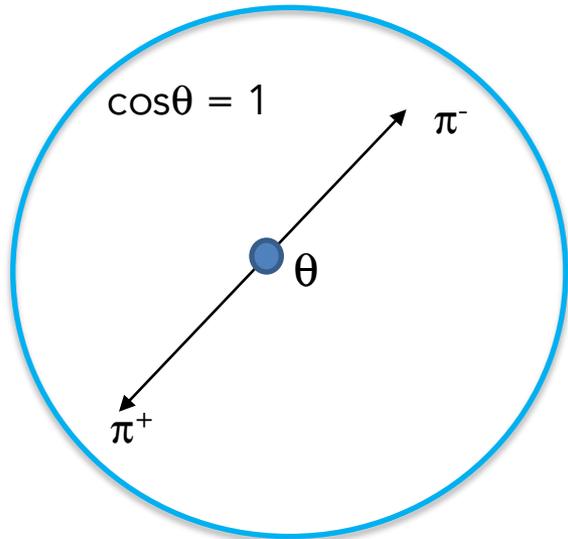


initial beam
 K_1 and K_2

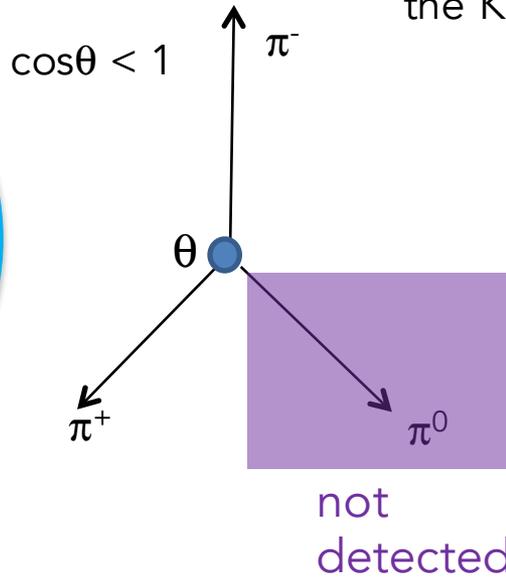
charged tracks
detector



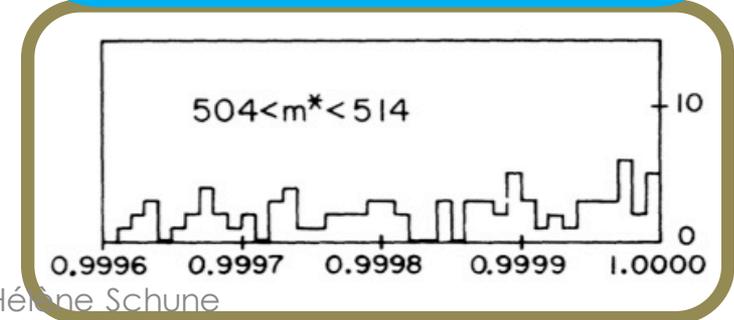
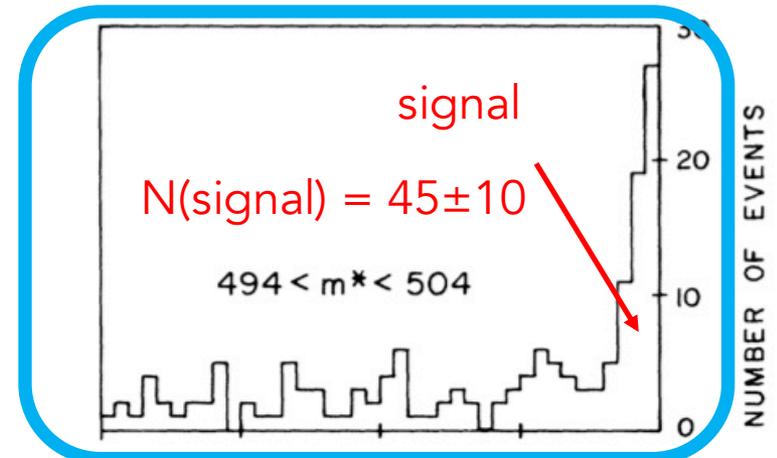
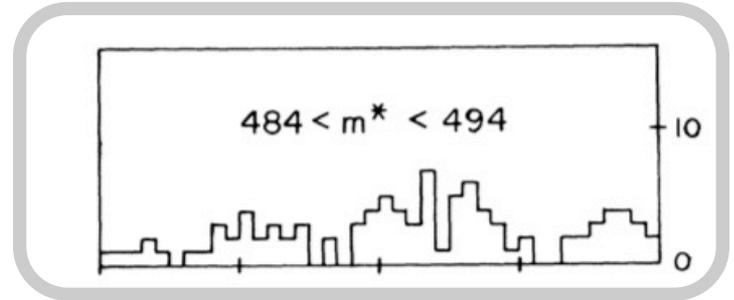
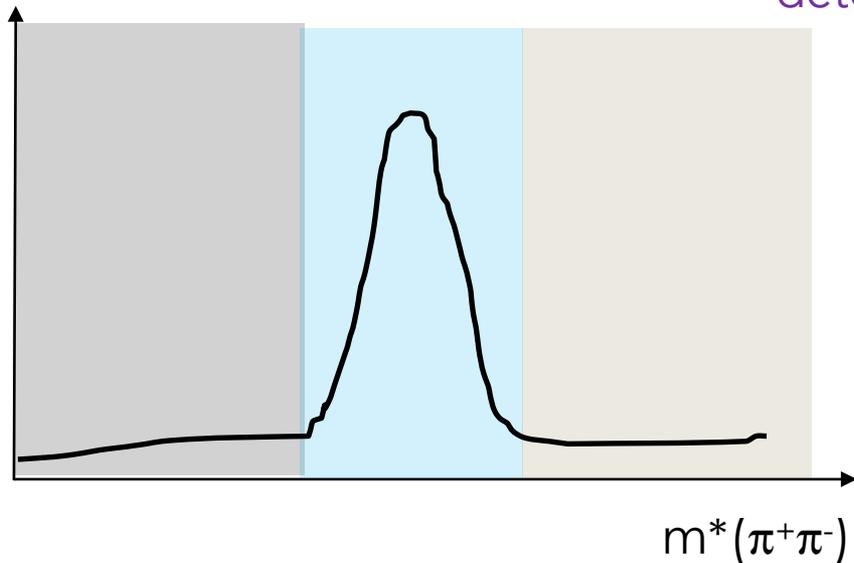
signal



background



- Two informations :
- The $\pi^+\pi^-$ invariant mass (m^*)
 - The opening angle between the two pions in the K center of mass frame



EVIDENCE FOR THE 2π DECAY OF THE K_2^0 MESON*†

J. H. Christenson, J. W. Cronin,[‡] V. L. Fitch,[‡] and R. Turley[§]
 Princeton University, Princeton, New Jersey
 (Received 10 July 1964)

1964

We would conclude therefore that K_2^0 decays to two pions with a branching ratio $R = (K_2^0 \rightarrow \pi^+ + \pi^-) / (K_2^0 \rightarrow \text{all charged modes}) = (2.0 \pm 0.4) \times 10^{-3}$ where the error is the standard deviation. As emphasized above, any alternate explanation of the effect requires highly nonphysical behavior of the three-body decays of the K_2^0 . The presence of a two-pion decay mode implies that **the K_2^0 meson is not a pure eigenstate of CP.** Expressed as

The Nobel Prize in Physics 1980



James Watson Cronin
 Prize share: 1/2



Val Logsdon Fitch
 Prize share: 1/2

The Nobel Prize in Physics 1980 was awarded jointly to James Watson Cronin and Val Logsdon Fitch "for the discovery of violations of fundamental symmetry principles in the decay of neutral K-mesons"

R. Turley was a PhD student
 J Christenson was a graduate student

« The discovery emphasizes, once again, that even almost self evident principles in science cannot be regarded fully valid until they have been critically examined in precise experiments. »

Today :

$$\frac{A(|K_2\rangle \rightarrow \pi\pi)}{A(|K_1\rangle \rightarrow \pi\pi)} = (2.271 \pm 0.017) 10^{-3} \quad 0.7 \% \text{ precision !}$$

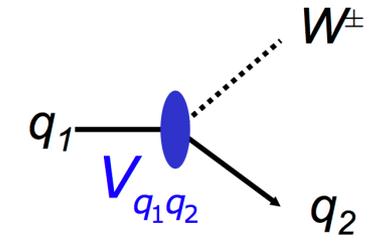
Experimental observation of CP violation in K decays
+ Cabibbo angle



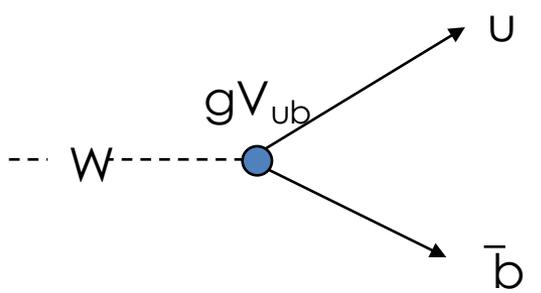
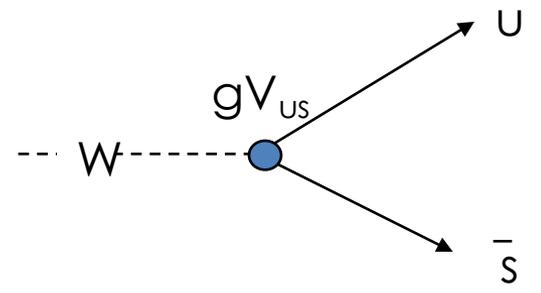
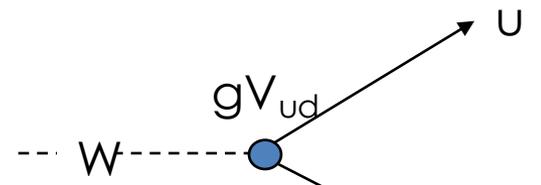
V_{CKM} Cabibbo-Kobayashi-Maskawa matrix

V_{CKM} Cabibbo-Kobayashi-Maskawa matrix

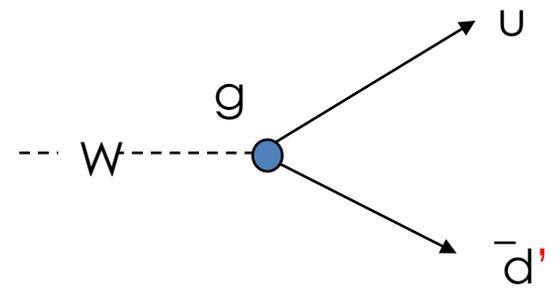
Two different way of seeing the charged interactions among quarks



In the basis dealing with mass eigenstates :



In the basis where : charged interactions are just between members of the same family and « CKM » is diagonal



Weak interaction eigenstates

\neq

Mass eigenstates (flavour or strong interaction eigenstates)

$$\begin{pmatrix} d' \\ s' \\ b' \end{pmatrix} = \begin{pmatrix} V_{ud} & V_{us} & V_{ub} \\ V_{cd} & V_{cs} & V_{cb} \\ V_{td} & V_{ts} & V_{tb} \end{pmatrix} \begin{pmatrix} d \\ s \\ b \end{pmatrix}$$

Two pink arrows point from the text above to the d' and b' components of the vector on the left side of the equation.

CP-Violation in the Renormalizable Theory of Weak Interaction

Makoto KOBAYASHI and Toshihide MASKAWA

Department of Physics, Kyoto University, Kyoto

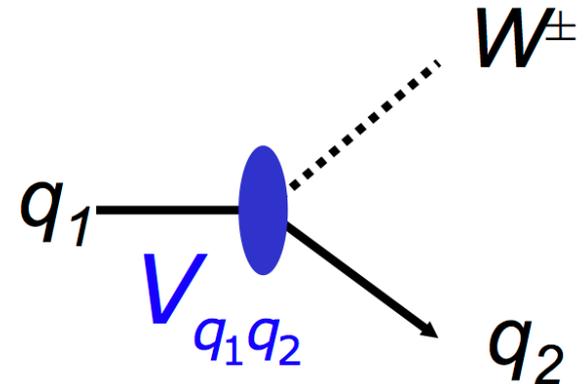
1973

Before the discovery of the 4th quark

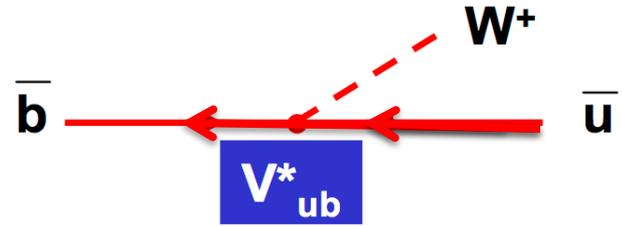
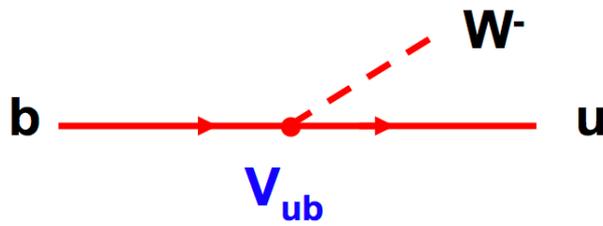
Prediction of the 3rd family

# families	# angles	# reducible phases	# irreducible phases
n	$n(n-1)/2$	$2n-1$	$n(n+1)/2 - (2n-1) = (n-1)(n-2)/2$
2	1		0
3	3		1
4	6		3

$$(u \quad c \quad t) \begin{pmatrix} V_{ud} & V_{us} & V_{ub} \\ V_{cd} & V_{cs} & V_{cb} \\ V_{td} & V_{ts} & V_{tb} \end{pmatrix} \begin{pmatrix} d \\ s \\ b \end{pmatrix}$$



V_{CKM} Cabibbo-Kobayashi-Maskawa matrix

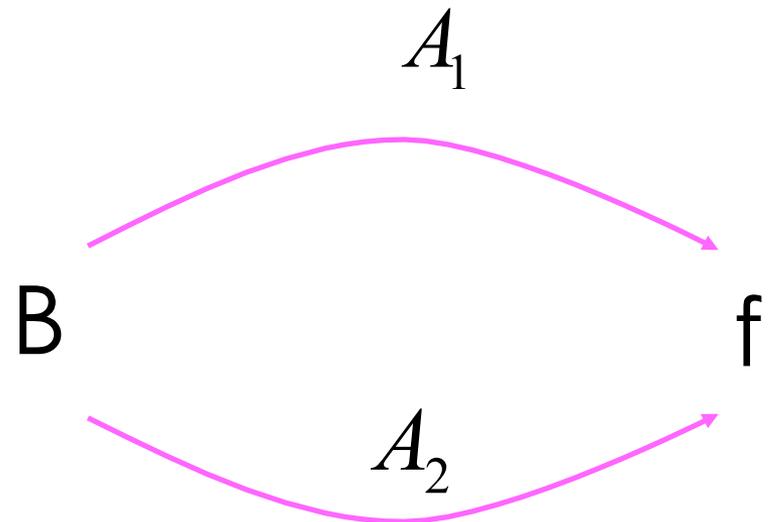
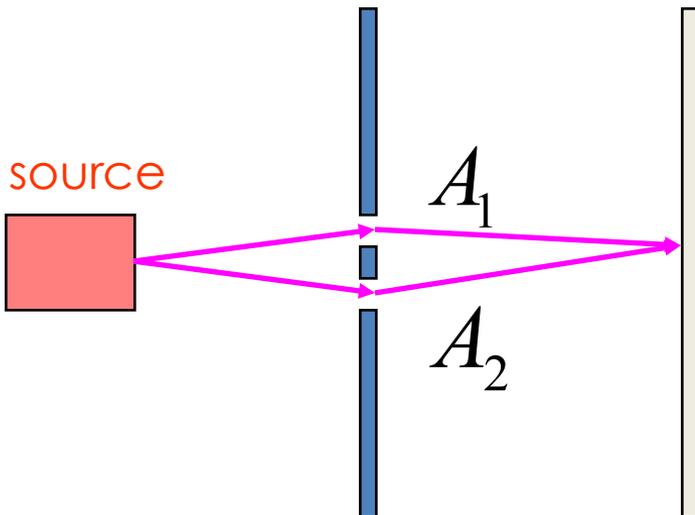


$V_{ub}^* \neq V_{ub} \rightarrow CP$ violation



CP

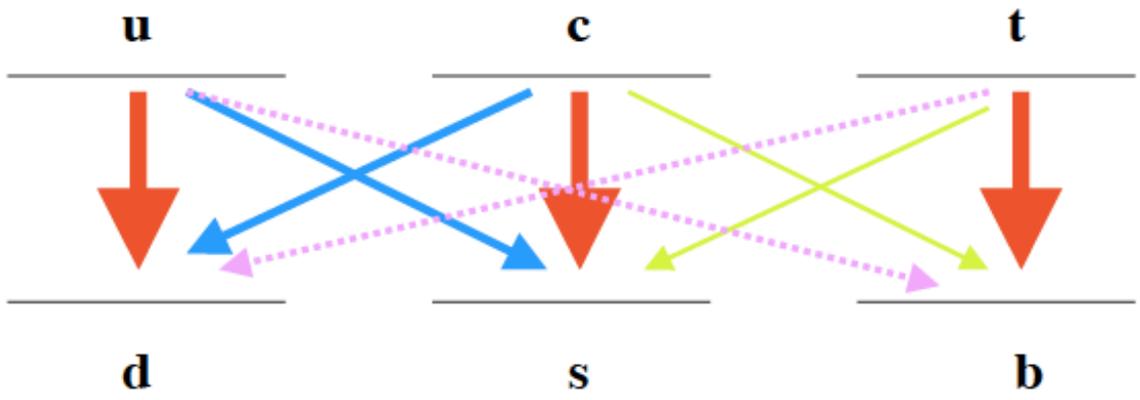
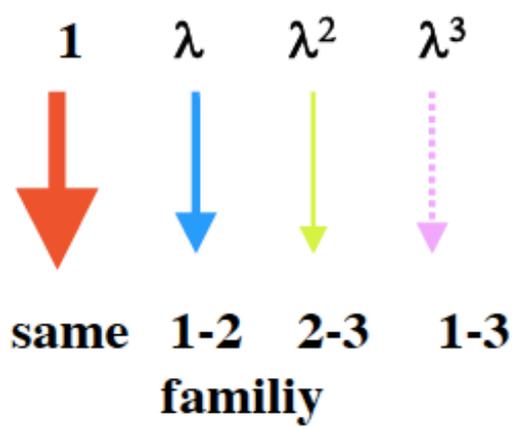
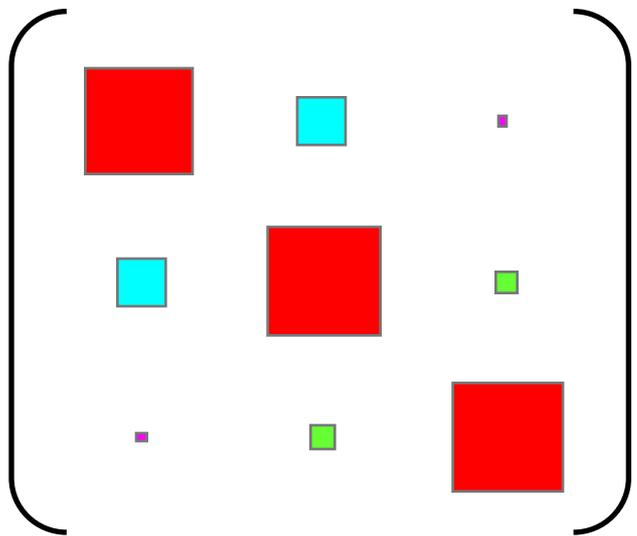
One amplitude : no sensitivity on phase ($|V_{ij}|^2 = |V_{ij}^*|^2$)



No prediction on the $V_{ij} \rightarrow$ they need to be measured

\rightarrow Experimental observations :

$$\begin{pmatrix} V_{ud} & V_{us} & V_{ub} \\ V_{cd} & V_{cs} & V_{cb} \\ V_{td} & V_{ts} & V_{tb} \end{pmatrix} =$$



Measuring triangles

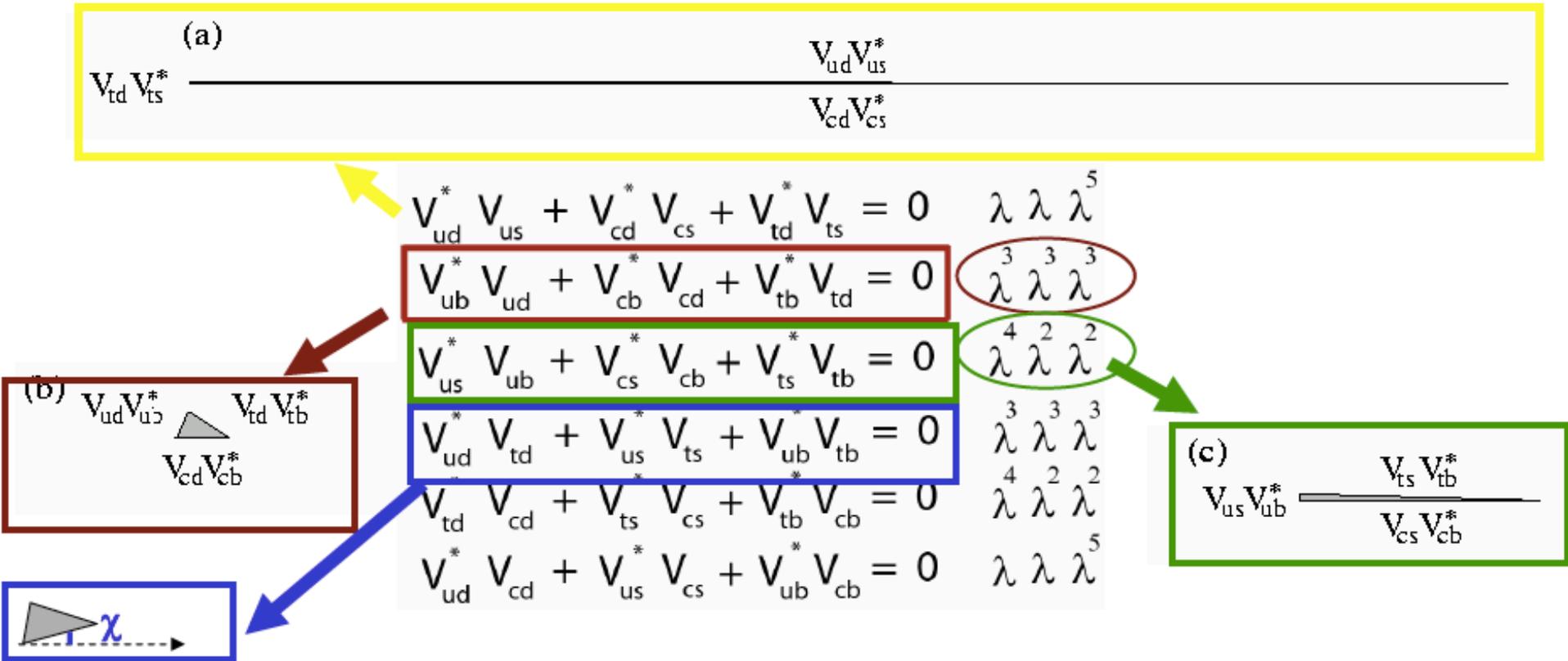
Stay within the 3 families

$$(u \quad c \quad t) \begin{pmatrix} V_{ud} & V_{us} & V_{ub} \\ V_{cd} & V_{cs} & V_{cb} \\ V_{td} & V_{ts} & V_{tb} \end{pmatrix} \begin{pmatrix} d \\ s \\ b \end{pmatrix}$$

Unitarity of V_{CKM} $VV^\dagger = V^\dagger V = \mathbf{1}$

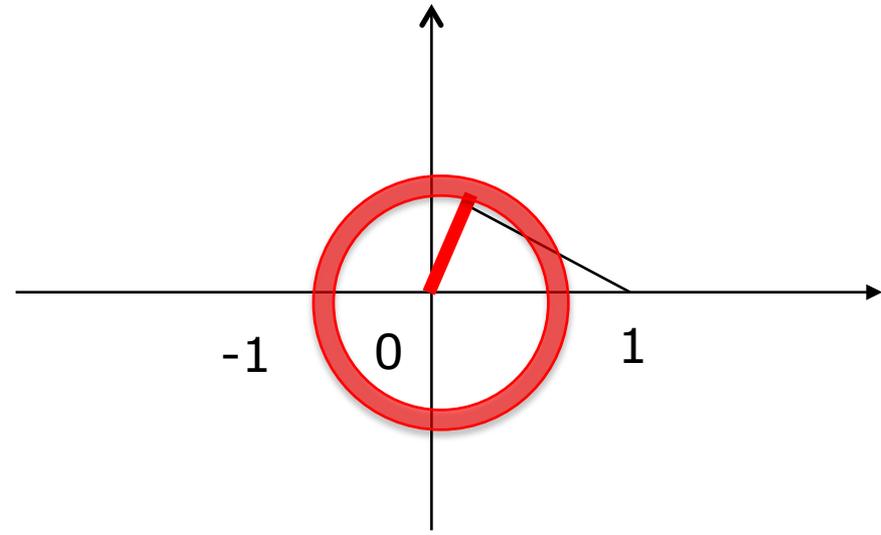
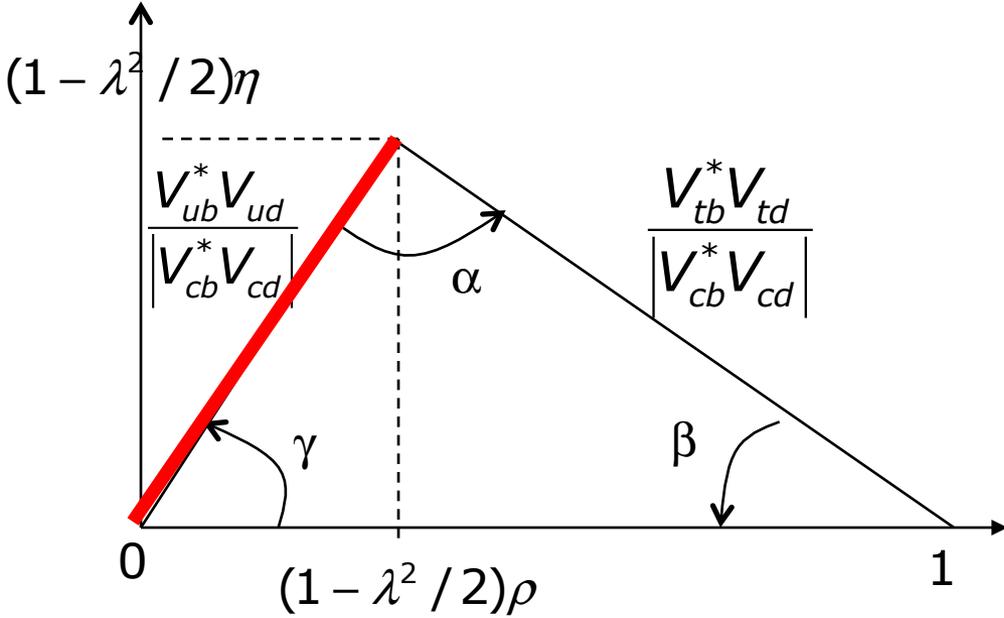
$\rightarrow 9$ relations $\sum_{k=1}^n V_{ik}V_{jk}^* = \delta_{ij},$

The non-diagonal elements of the matrix products correspond to 6 triangle equations

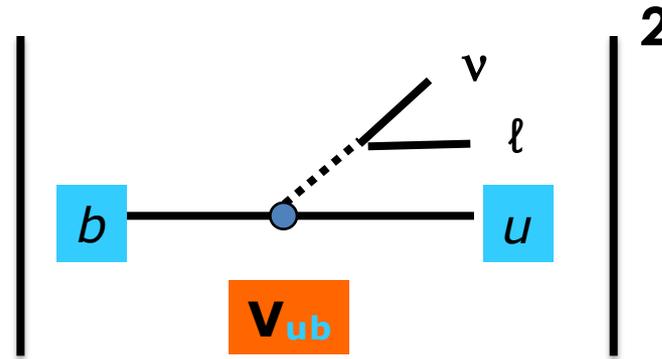
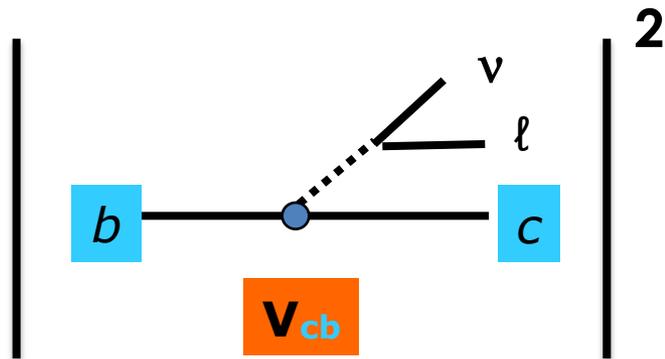


They all have the same area, proportionnal to the amount of CP violation in the SM

Measurements of the unitarity triangle parameters : some examples

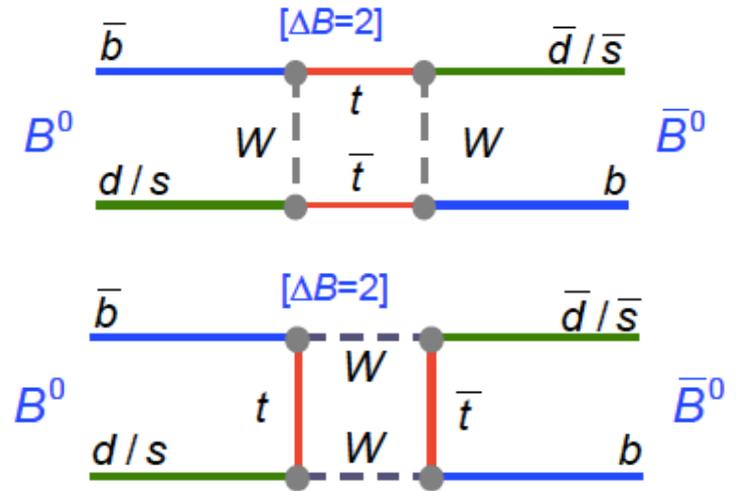
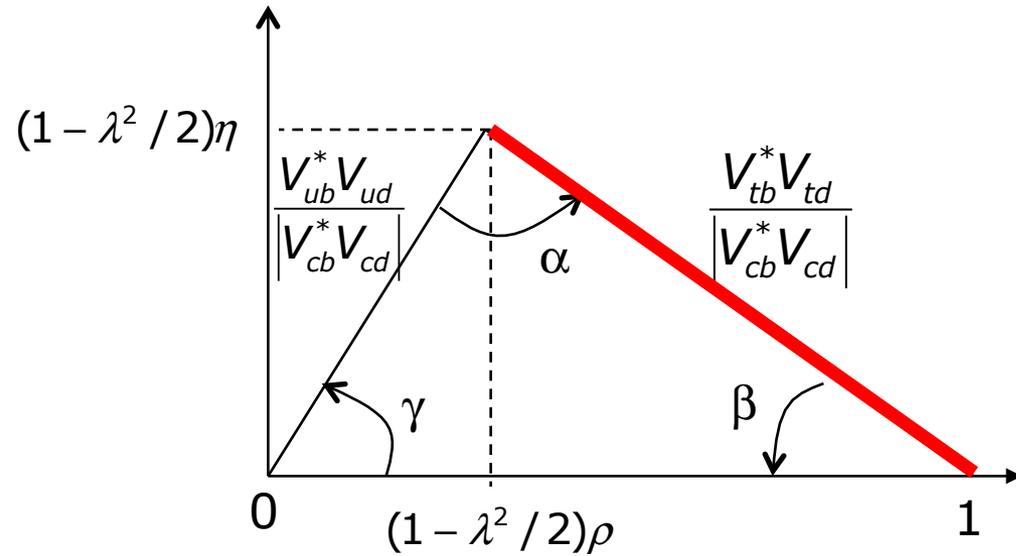


Rates of semileptonic B decays



Conceptually simple, complicated by QCD

The other side : $B^0-\bar{B}^0$ oscillations



Diagrams involving V_{td} or V_{ts}



The mixing phenomenon

Pairs of self-conjugate mesons that can be transformed to each other via flavour changing weak interaction transitions are:

$$|K^0\rangle = |\bar{s}d\rangle \quad |D^0\rangle = |c\bar{u}\rangle \quad |B_d^0\rangle = |\bar{b}d\rangle \quad |B_s^0\rangle = |\bar{b}s\rangle$$

They are **flavour eigenstates** with definite quark content

- useful to understand particle production and decay

$$|B^0\rangle, |\bar{B}^0\rangle$$

Apart from the flavour eigenstates there are **mass eigenstates**:

- eigenstates of the Hamiltonian
- states of definite mass and lifetime
- They are propagating through space-time

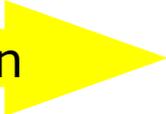
$$|B_L\rangle, |B_H\rangle$$

$$|B_L\rangle = p|B^0\rangle + q|\bar{B}^0\rangle$$

$$|B_H\rangle = p|B^0\rangle - q|\bar{B}^0\rangle$$

Since flavour eigenstates are not mass eigenstates, the flavour eigenstates are mixed with one another as they propagate through space and time

$$|B_{H,L}(t)\rangle = e^{-i\left(M_{H,L} - i\frac{\Gamma_{H,L}}{2}\right)t} |B_{H,L}(t=0)\rangle + \begin{aligned} |B_L\rangle &= p |B^0\rangle + q |\bar{B}^0\rangle \\ |B_H\rangle &= p |B^0\rangle - q |\bar{B}^0\rangle \end{aligned}$$

Time evolution 

The probability to observe a B^0 at time t if a B^0 was produced at time $t=0$ is :

$$\left| \langle B^0 | H | B^0(t) \rangle \right|^2 = \frac{e^{-\Gamma t}}{2} (1 + \cos \Delta m t)$$

The probability to observe a \bar{B}^0 at time t if a \bar{B}^0 was produced at time $t=0$ is :

$$\left| \langle \bar{B}^0 | H | B^0(t) \rangle \right|^2 = \frac{e^{-\Gamma t}}{2} (1 - \cos \Delta m t)$$

This is the mixing phenomenon !

$$\frac{N_{Unmixed} - N_{Mixed}}{N_{Unmixed} + N_{Mixed}} \sim \cos \Delta m t$$

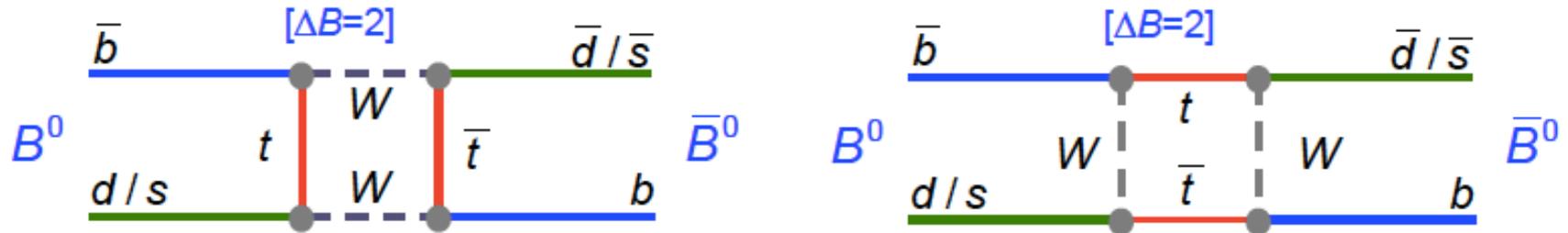
Simplified formulae assuming that the two mass eigenstates have the same lifetime and neglecting CP violation ($q/p=1$)



Let's come back to the unitarity triangle

Δm can be computed in the Standard Model

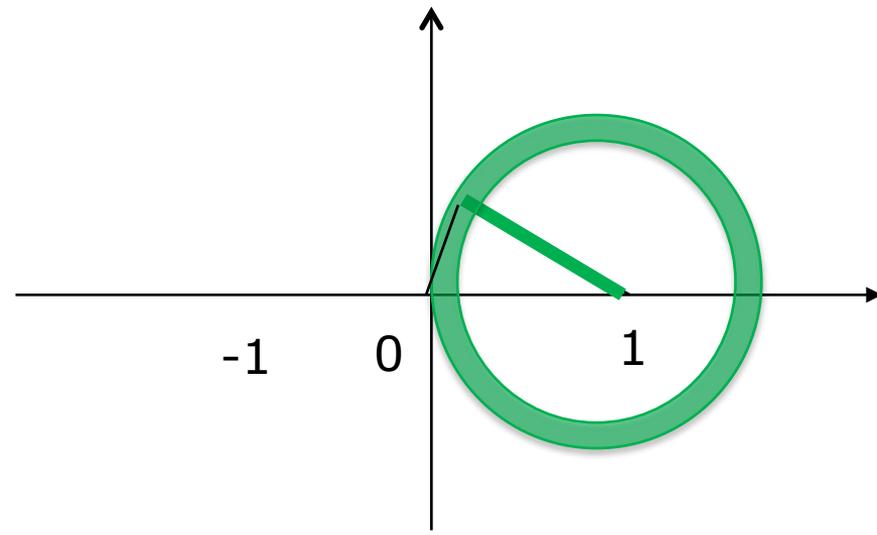
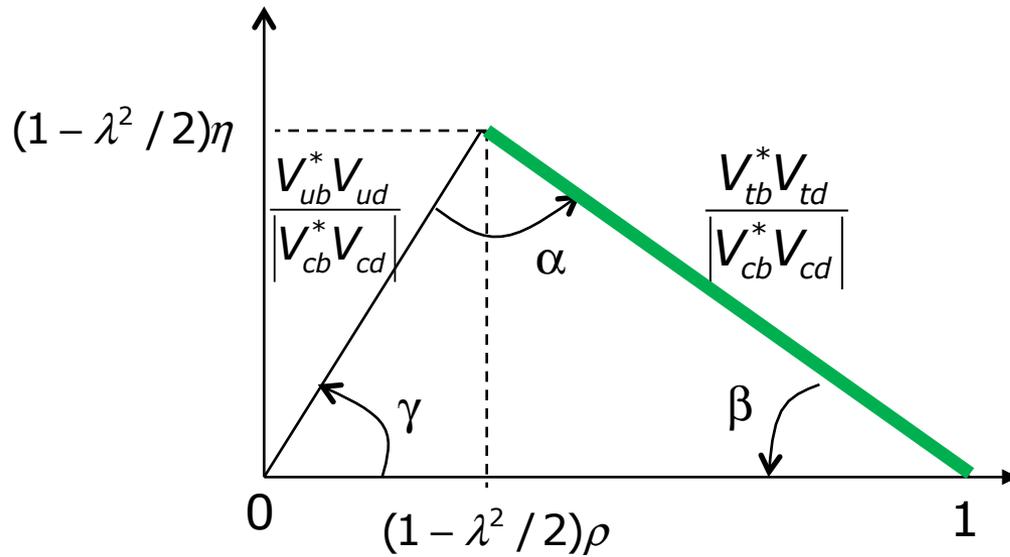
Effective FCNC Processes (CP conserving — **top** loop dominates in box diagram):



$$\Delta m_q = \frac{G_F^2}{6\pi^2} m_{B_q} m_W^2 \eta_B S(x_t) f_{B_q}^2 B_q |V_{tq} V_{tb}^*|^2 \quad (\text{for } q = d, s)$$

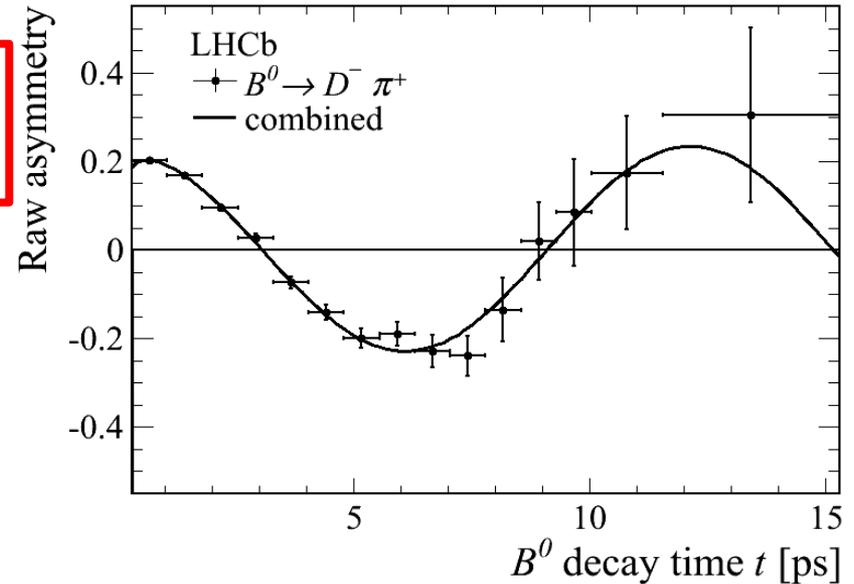
Perturbative QCD → η_B
CKM Matrix Elements → $|V_{tq} V_{tb}^*|^2$
Loop integral (top loop dominates) → $S(x_t)$
Non-perturbative QCD : dominant theoretical uncertainty → $f_{B_q}^2 B_q$

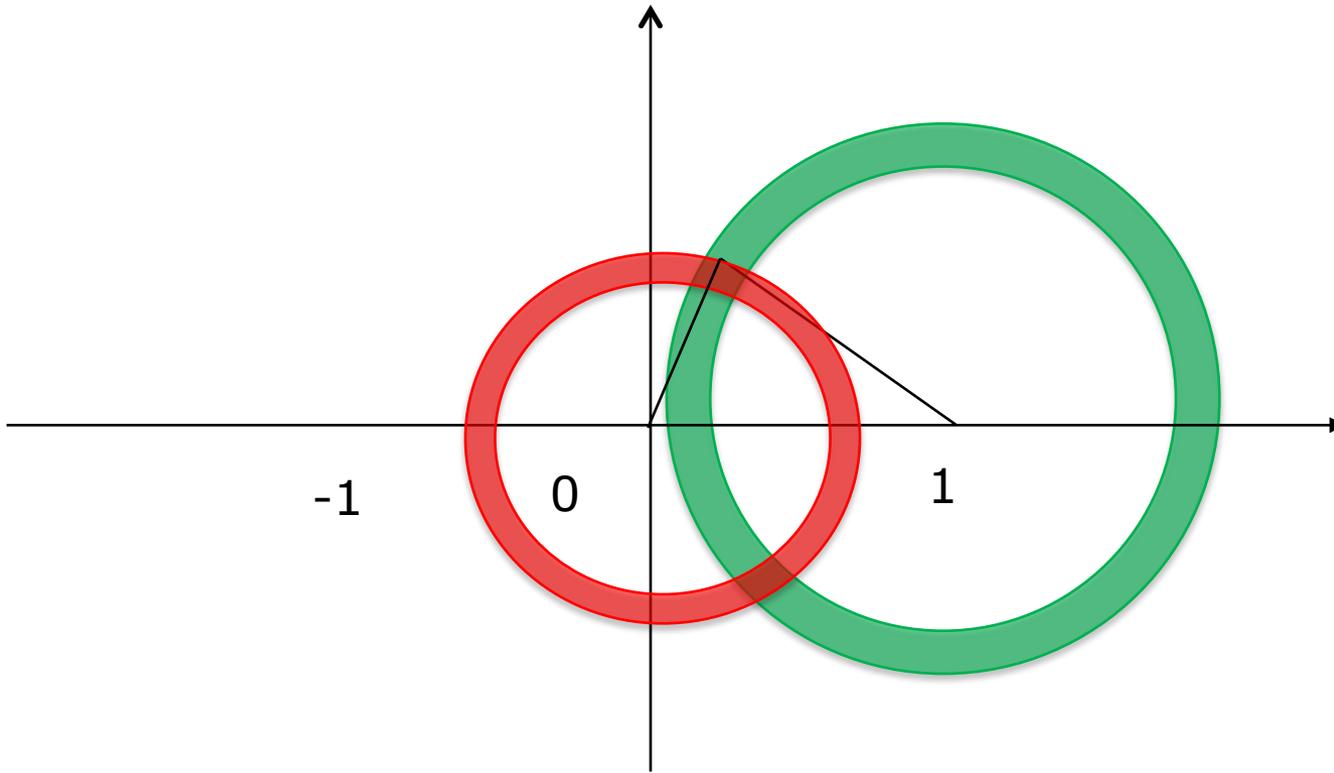
The other side : $B^0-\bar{B}^0$ oscillations



$$\frac{N_{Unmixed} - N_{Mixed}}{N_{Unmixed} + N_{Mixed}} \sim \cos \Delta mt$$

$$\Delta m_d \propto |V_{td} V_{tb}^*|^2$$

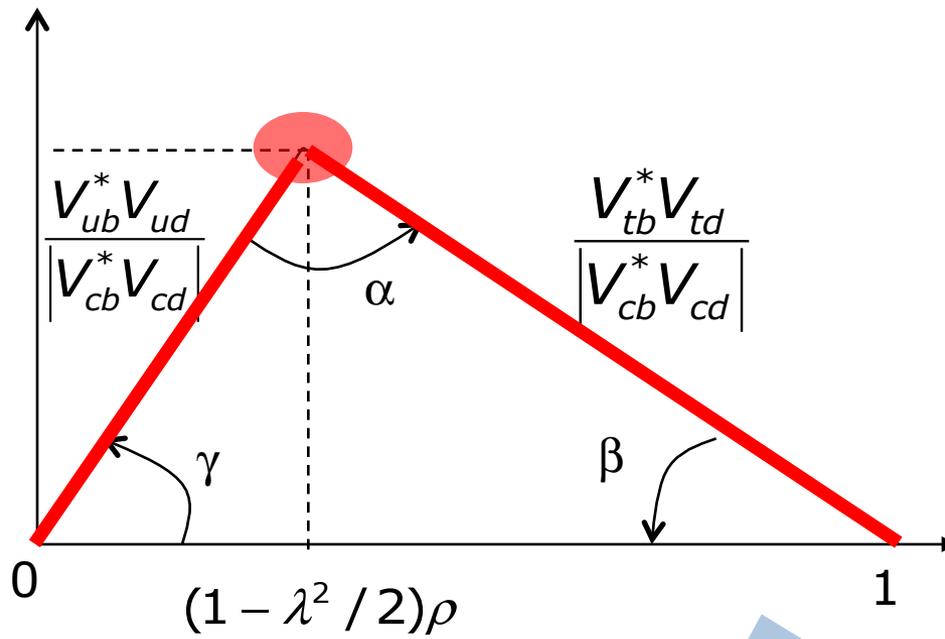




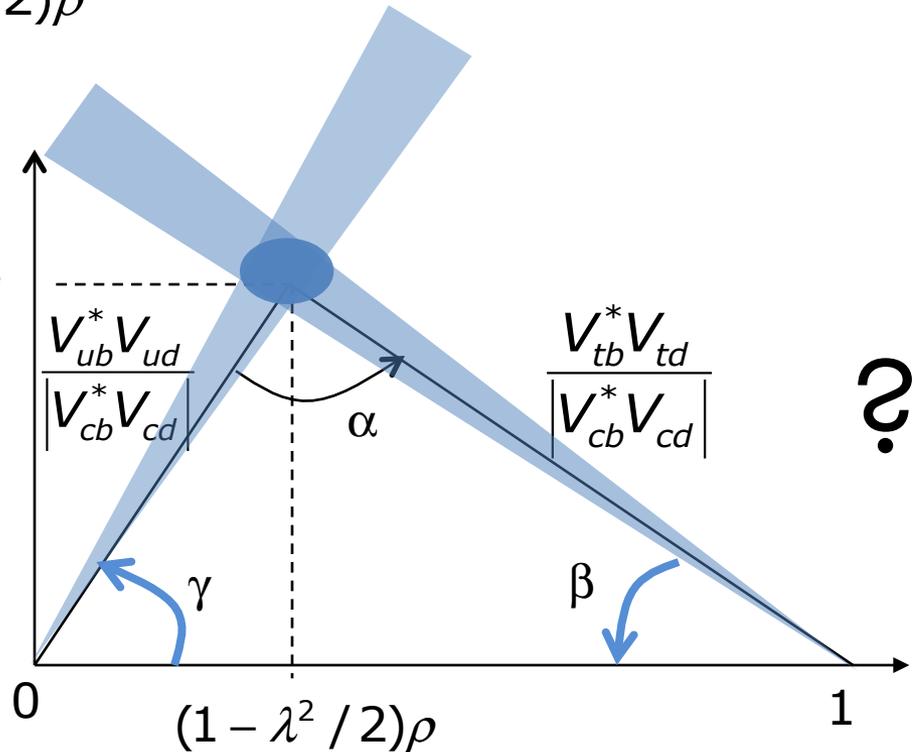
Are the two types of measurements compatible ?

Is

$$(1 - \lambda^2 / 2)\eta$$



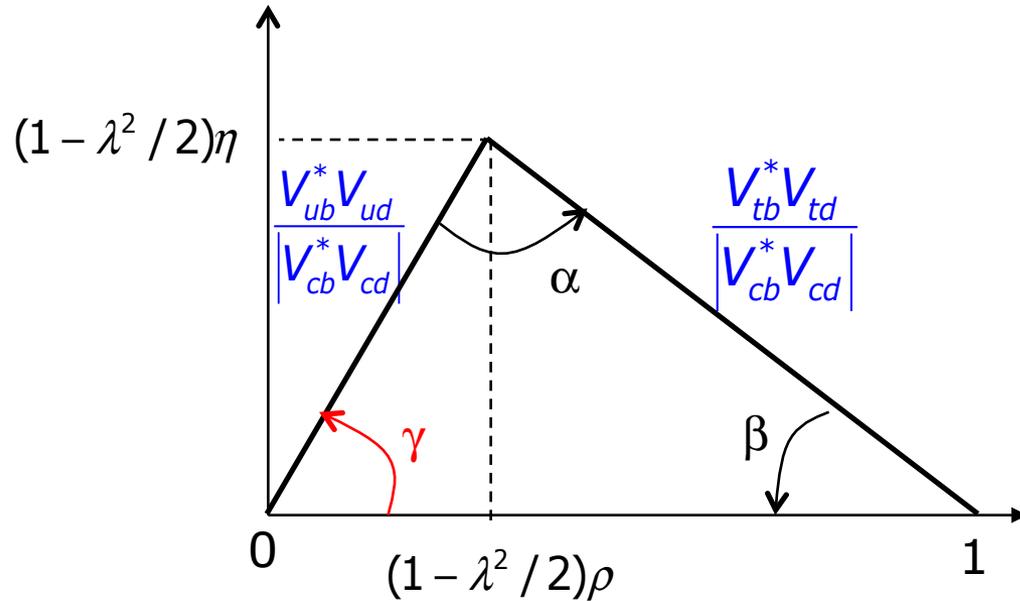
in
agreement
with



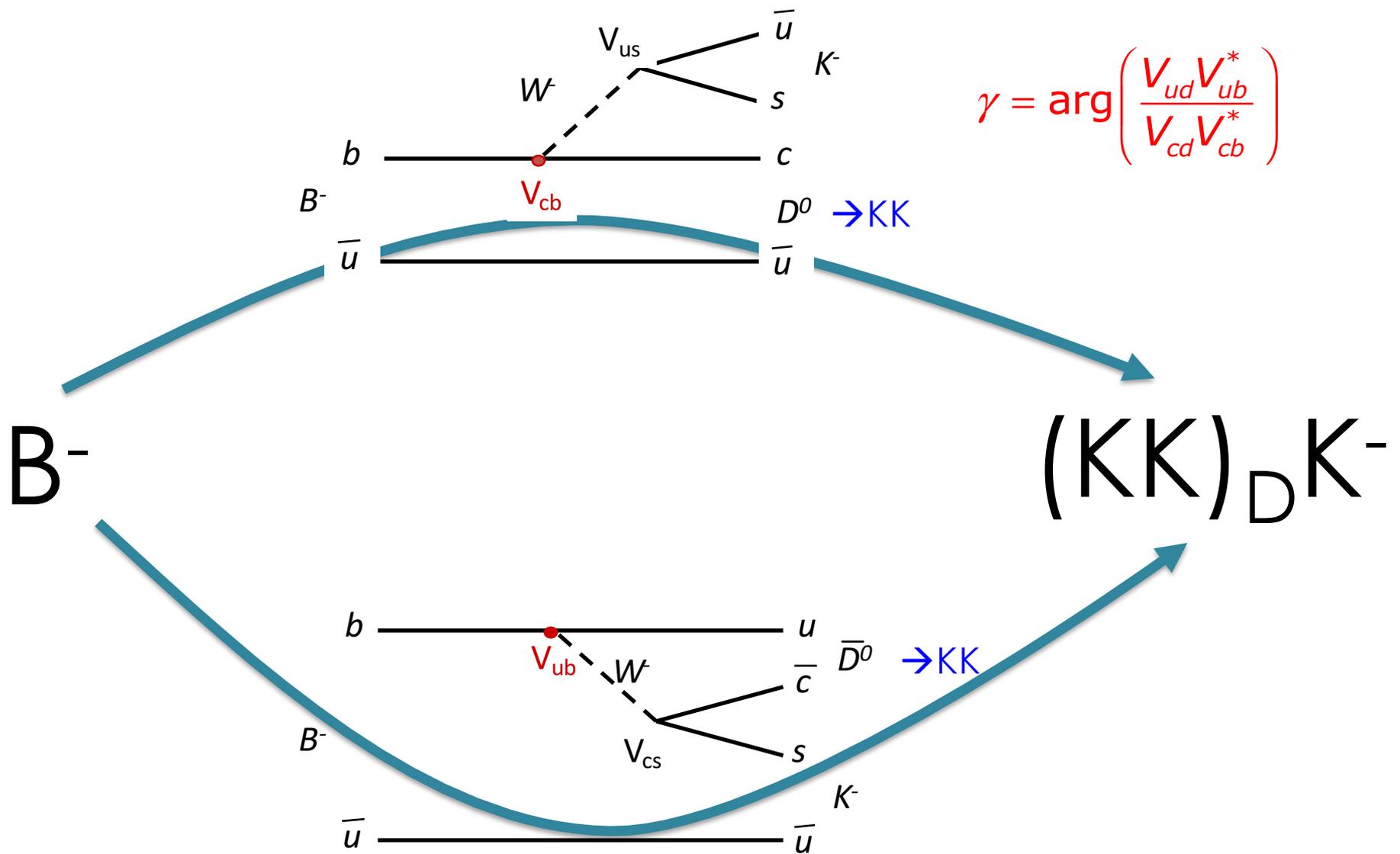
“the” unitarity triangle :

$$V_{ub}^* V_{ud} + V_{cb}^* V_{cd} + V_{tb}^* V_{td} = 0$$

CP violation measurement



$$\gamma = \arg \left(\frac{V_{ud} V_{ub}^*}{V_{cd} V_{cb}^*} \right)$$

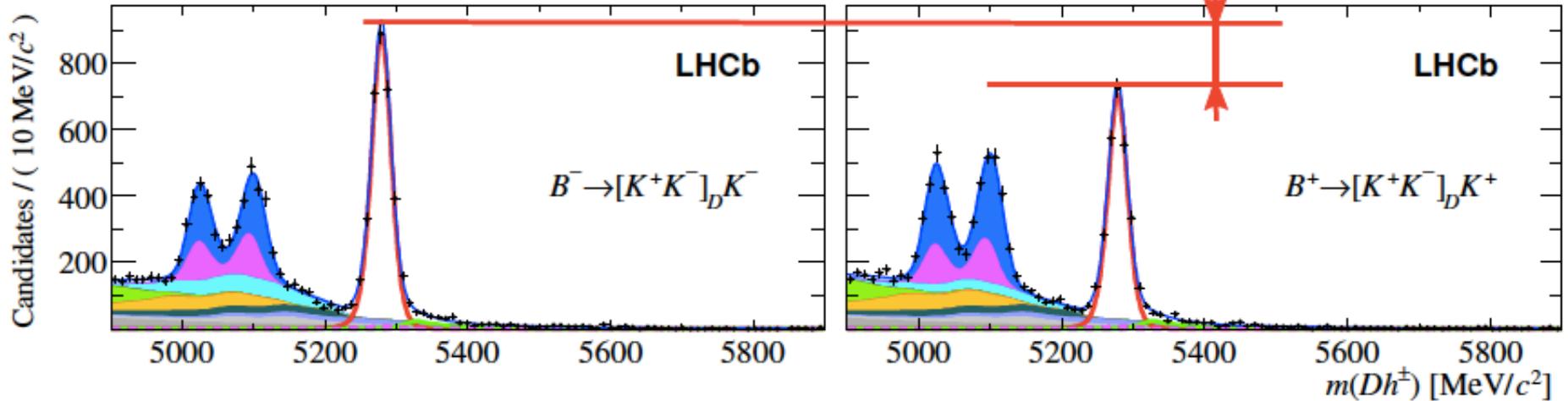


$(KK)_D K^\pm$

Start with the same amount of B^+ and B^-
 Count $N(B^+ \rightarrow (KK)_D K^+)$ and $N(B^- \rightarrow (KK)_D K^-)$

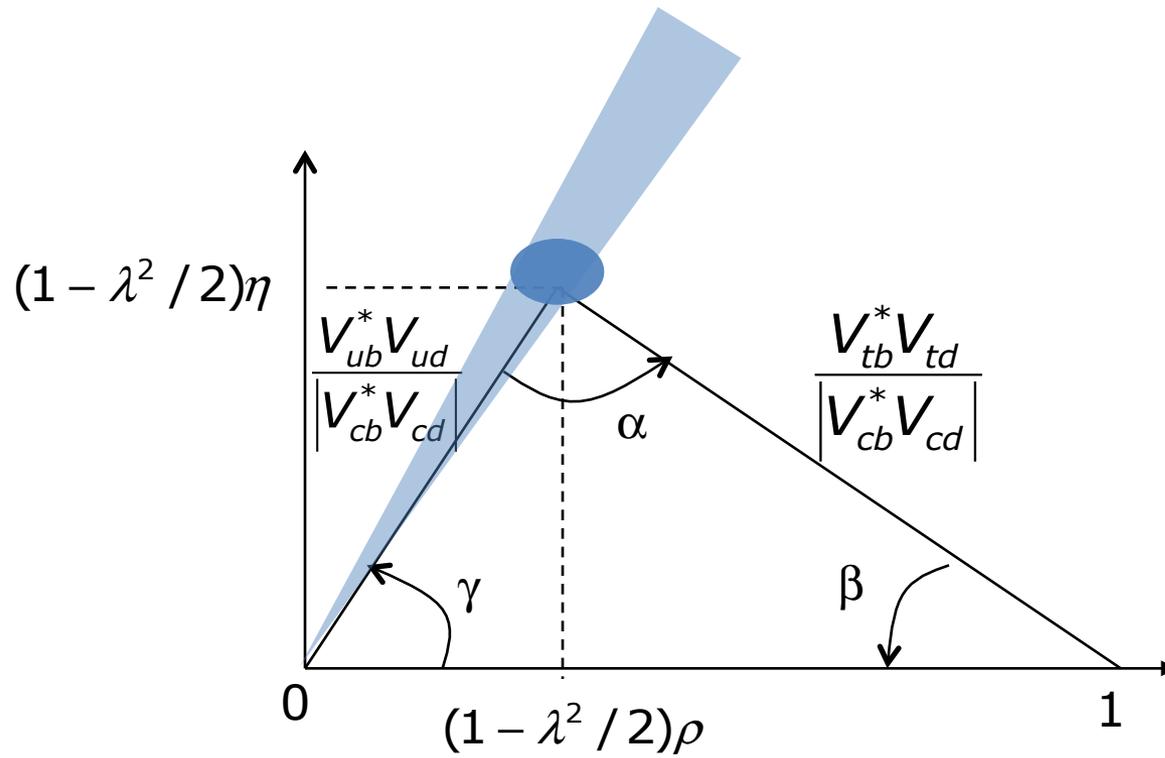


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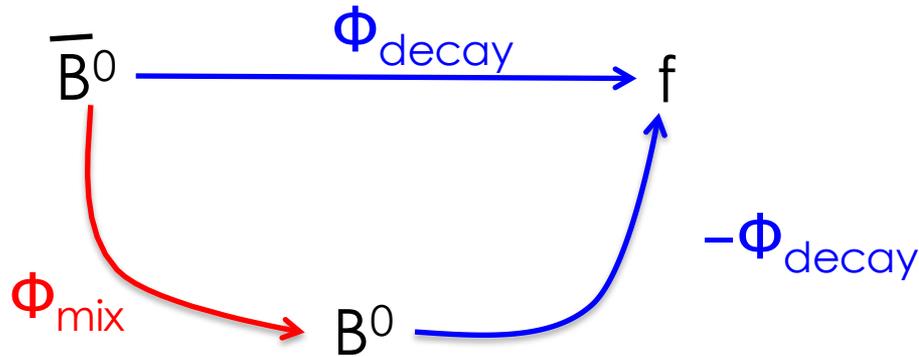


$$A_K^{KK} = +0.126 \pm 0.014 \text{ (stat)} \pm 0.002 \text{ (syst)}$$

significantly different from 0 !

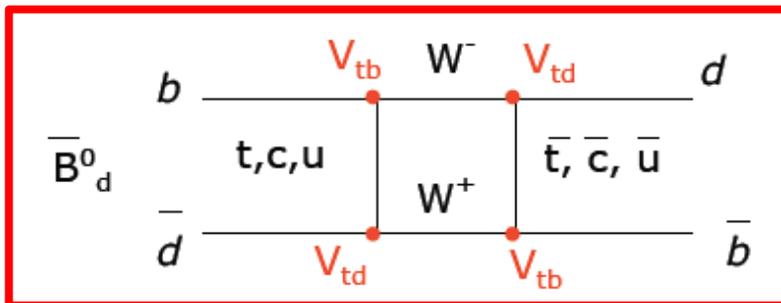


An example of CP induced by the interference between mixing and decay : the β angle

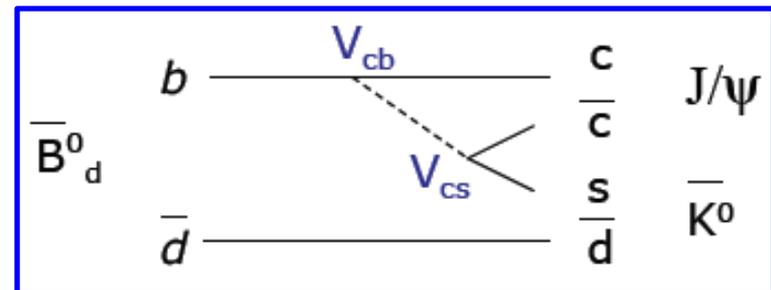


$$\Phi_d = \Phi_{mix} - 2 \Phi_{decay}$$

Mixing



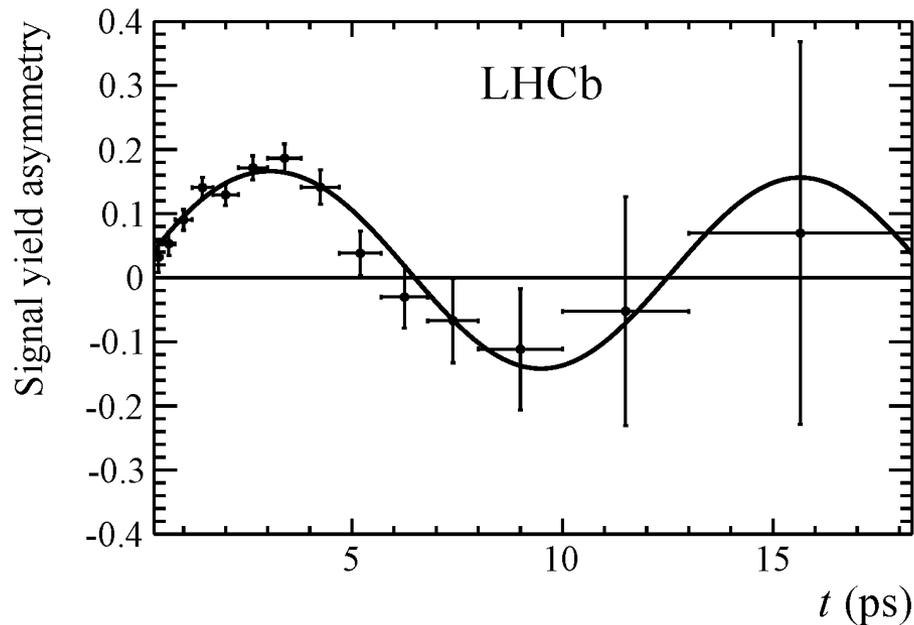
Decay



$$a_{f_{CP}}(t) = \frac{\text{Prob}(\overline{B^0}(t) \rightarrow f_{CP}) - \text{Prob}(B^0(t) \rightarrow f_{CP})}{\text{Prob}(\overline{B^0}(t) \rightarrow f_{CP}) + \text{Prob}(B^0(t) \rightarrow f_{CP})} =$$

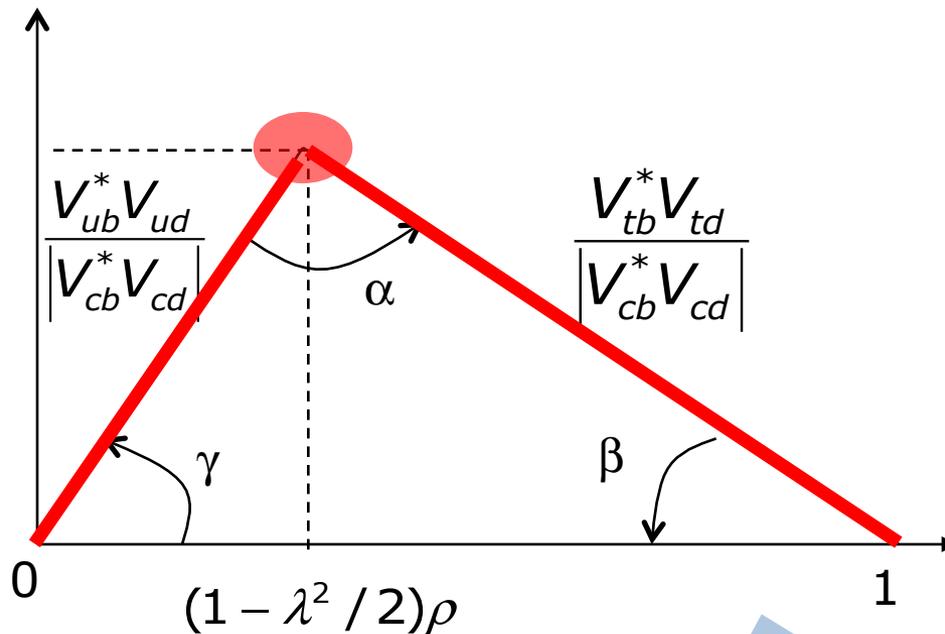
$$= \sin(2\beta) \sin(\Delta m t)$$

Pionnered by the B-factories

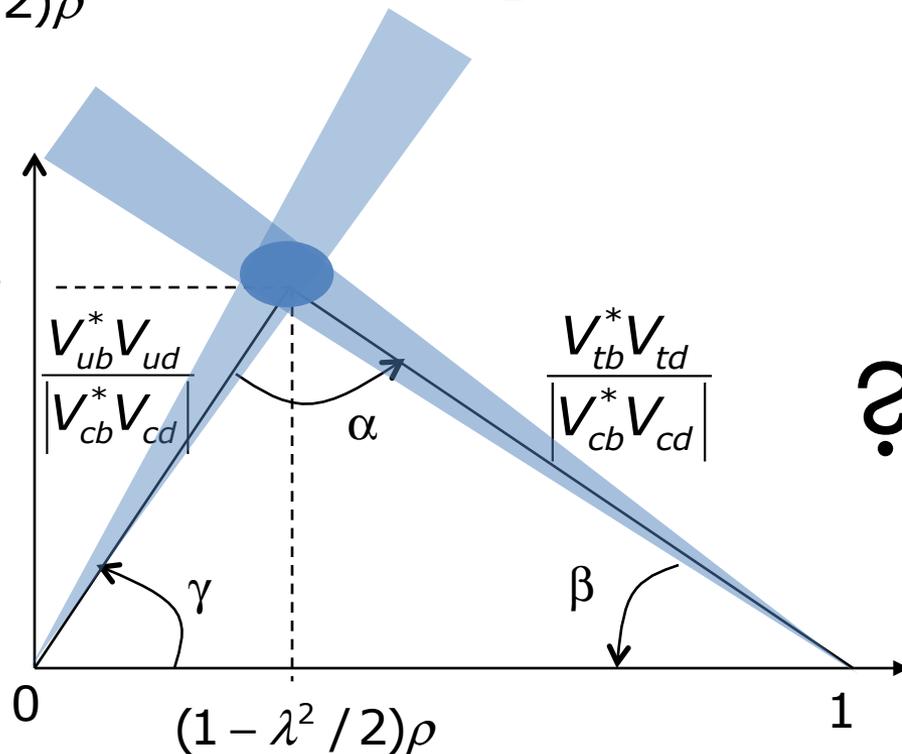


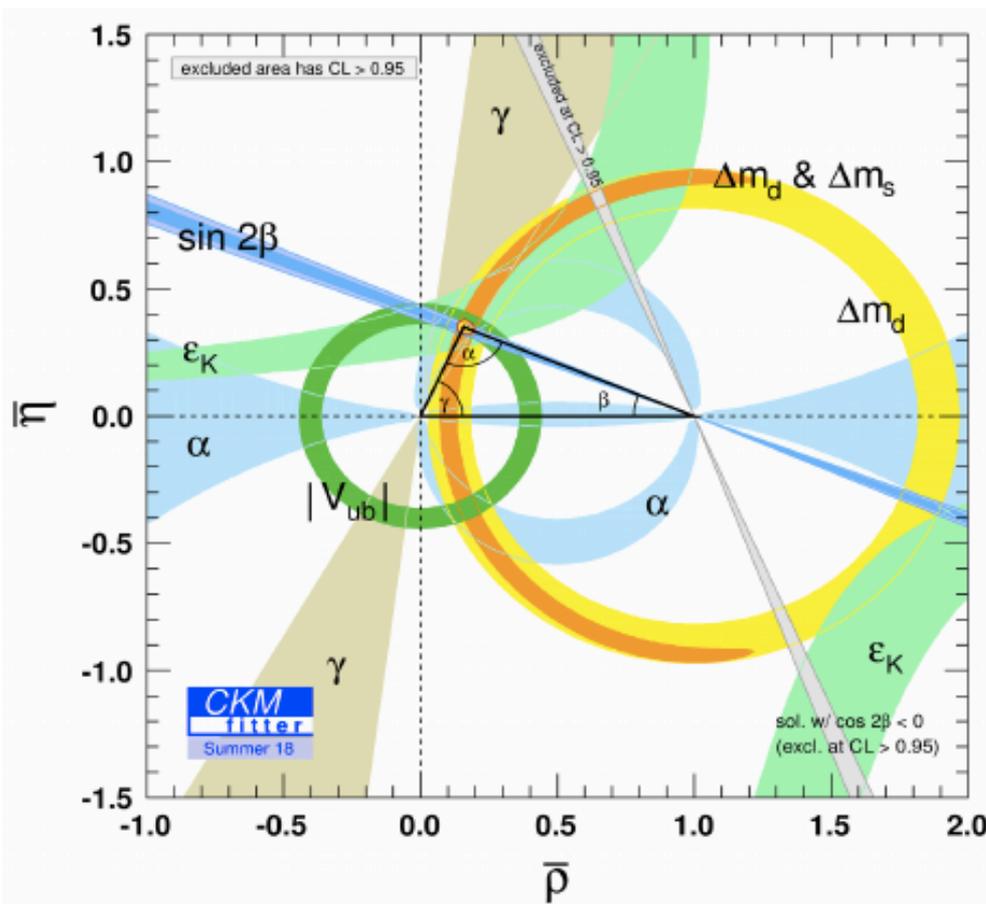
Is

$$(1 - \lambda^2 / 2)\eta$$



in
agreement
with





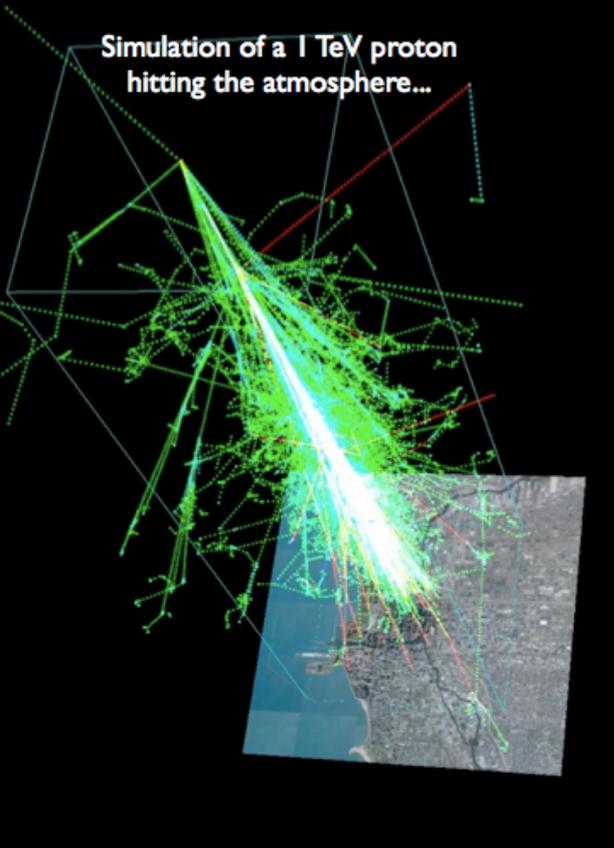
$$\bar{\rho} = 0.1577^{+0.0096}_{-0.0074} \text{ (5\% unc.)}$$

$$\bar{\eta} = 0.3493^{+0.0095}_{-0.0071} \text{ (2\% unc.)}$$

Sides and angles measurements in good agreement

The CKM model of CP violation has been confirmed

At the electroweak scale, the CKM mechanism dominates CP Violation



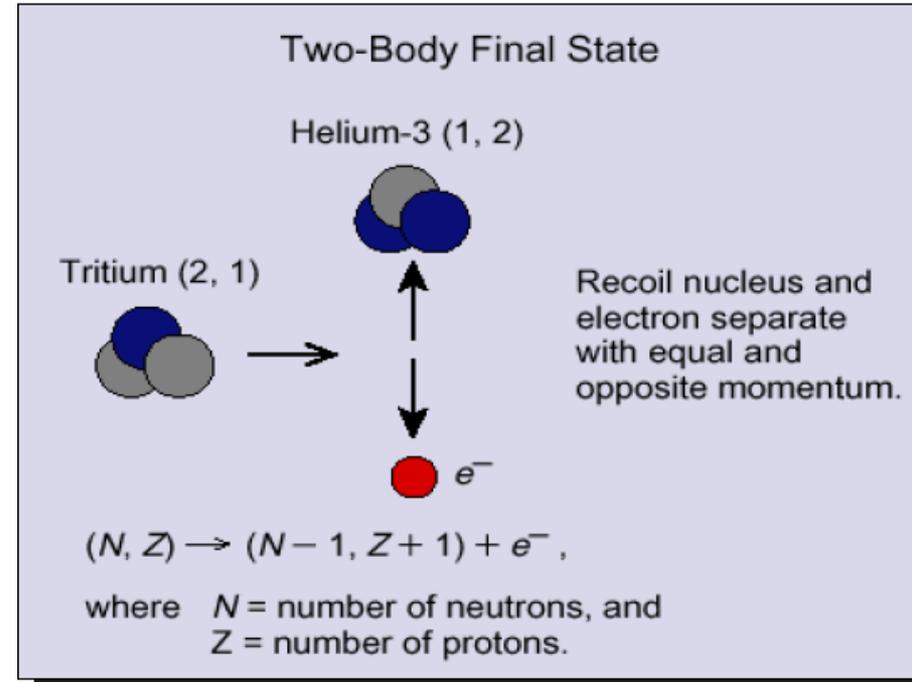
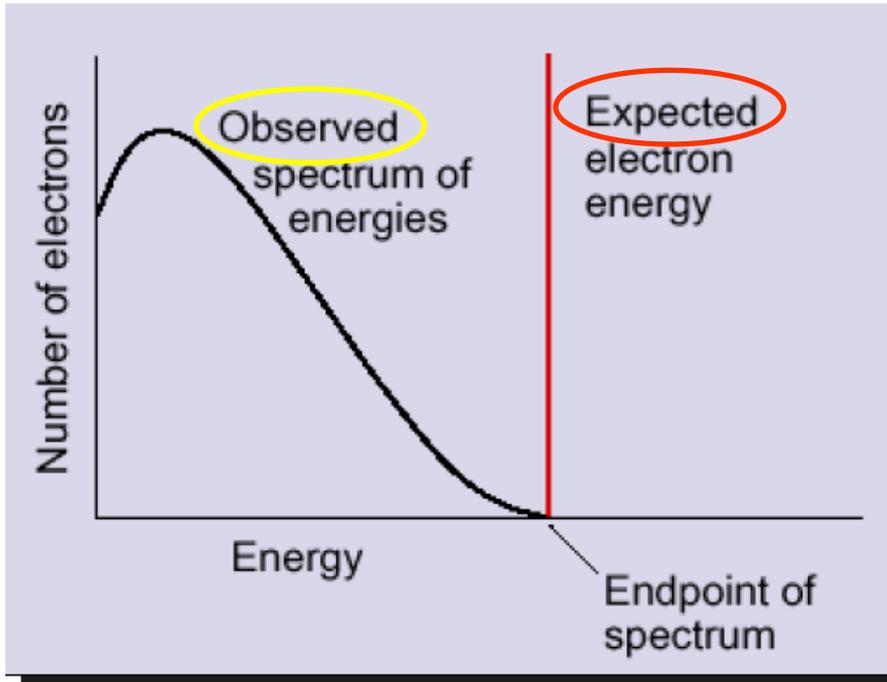
- Anti-matter in cosmic rays
 - No sign of light emission (anti-galaxy ...)
 - No significant sign of anti-nuclei (anti-He⁴ ...)
- Searches on-going



But the CP violation phase of the SM is orders of magnitude too small

Neutrinos

The β decay



If β decay proceeds through $n \rightarrow p e^-$ the energy conservation predicts a monochromatic spectrum

1914 Chadwick observes a continuous spectrum ...

→ Energy conservation is violated (Bohr) or an other particle is in the game (Pauli) :
 β decay : $n \rightarrow p e^- \bar{\nu}_e$

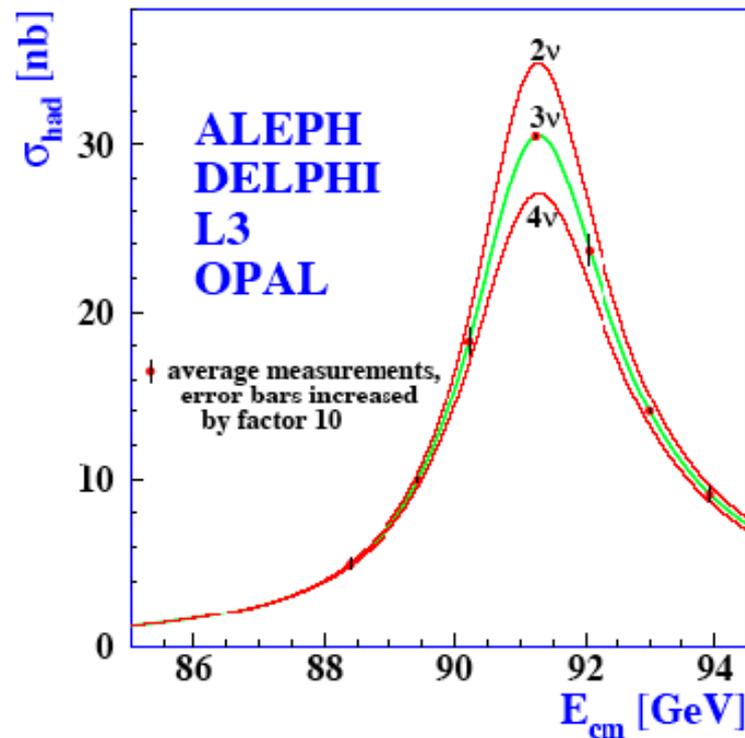
Fermionic sector of the SM :

	I	II	III
mass	$\approx 2.4 \text{ MeV}/c^2$	$\approx 1.275 \text{ GeV}/c^2$	$\approx 172.44 \text{ GeV}/c^2$
charge	$2/3$	$2/3$	$2/3$
spin	$1/2$	$1/2$	$1/2$
QUARKS	u up	c charm	t top
	$\approx 4.8 \text{ MeV}/c^2$	$\approx 95 \text{ MeV}/c^2$	$\approx 4.18 \text{ GeV}/c^2$
	$-1/3$	$-1/3$	$-1/3$
	$1/2$	$1/2$	$1/2$
	d down	s strange	b bottom
	$\approx 0.511 \text{ MeV}/c^2$	$\approx 105.67 \text{ MeV}/c^2$	$\approx 1.7768 \text{ GeV}/c^2$
	-1	-1	-1
	$1/2$	$1/2$	$1/2$
	e electron	μ muon	τ tau
LEPTONS	$< 2.2 \text{ eV}/c^2$	$< 1.7 \text{ MeV}/c^2$	$< 15.5 \text{ MeV}/c^2$
	0	0	0
	$1/2$	$1/2$	$1/2$
	ν_e electron neutrino	ν_μ muon neutrino	ν_τ tau neutrino

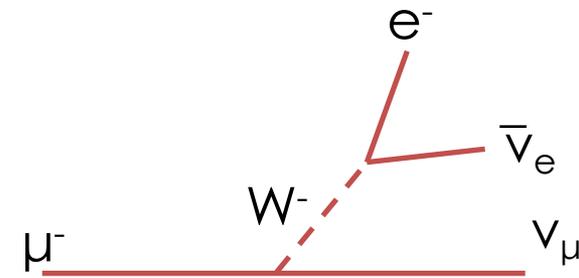
Lepton number conservation

3 types of neutrinos with $m = 0$

$$N_\nu = 2.9840 \pm 0.0082$$

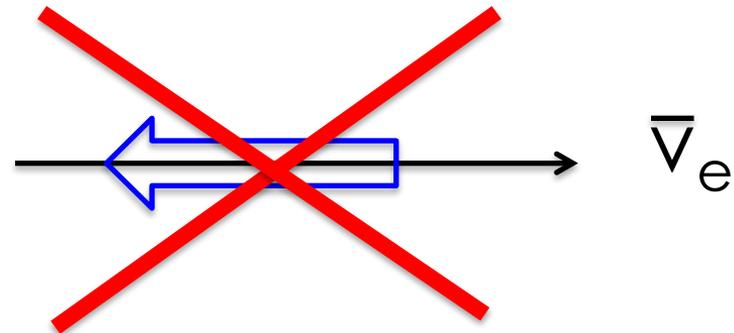
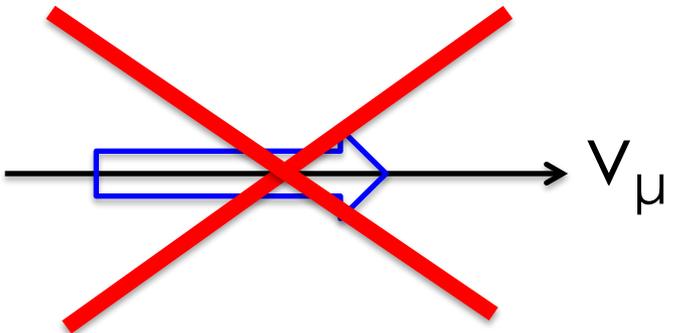
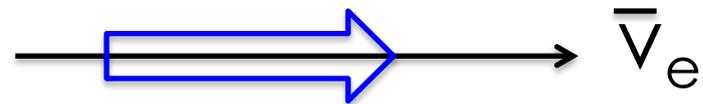
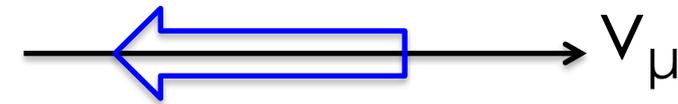


matrix element for the Feynman graph:



$$M = \left(\frac{g}{\sqrt{2}} \bar{u}_{\nu_\mu} \gamma^\mu \frac{1}{2} (1 - \gamma_5) u_\mu \right) \frac{1}{M_W^2 - q^2} \left(\frac{g}{\sqrt{2}} \bar{u}_e \gamma_\mu \frac{1}{2} (1 - \gamma_5) u_{\nu_e} \right)$$

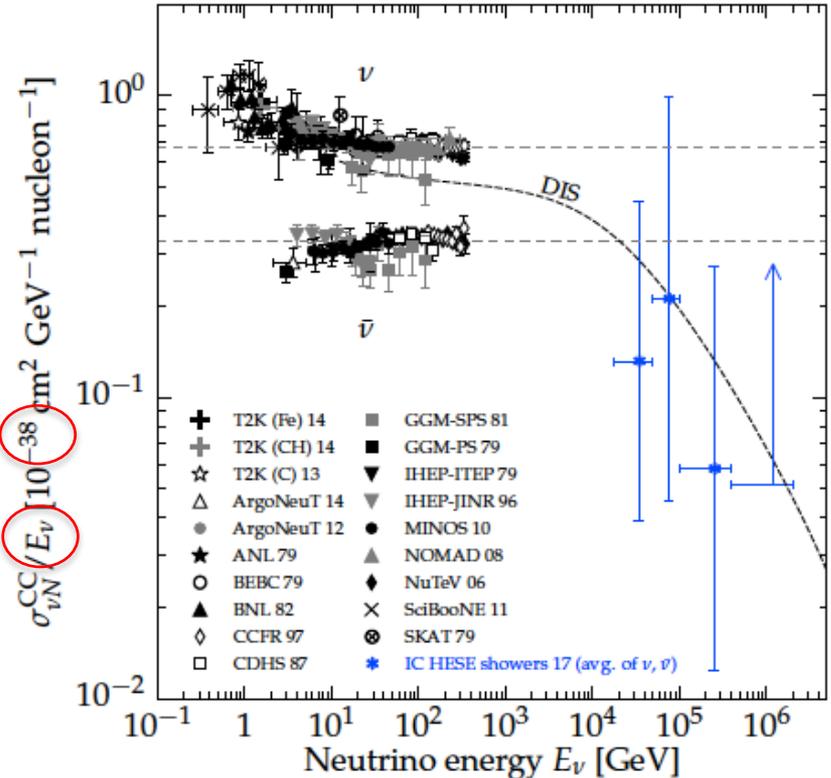
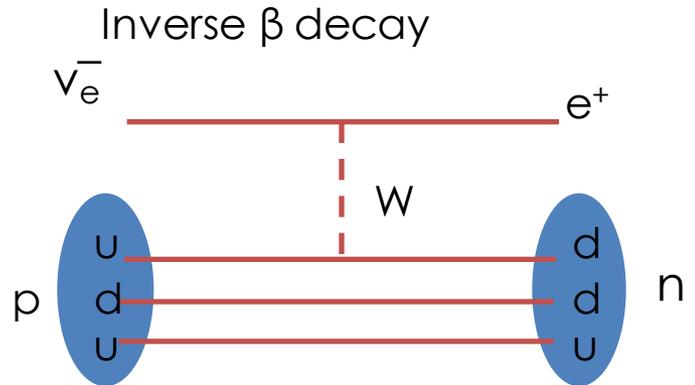
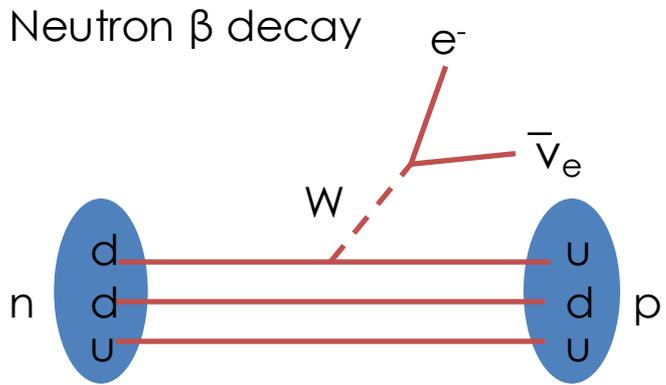
$$G_F = \frac{\sqrt{2}}{8} \frac{g^2}{M_W^2}$$



The ν is left handed (the anti-neutrino is right handed)

How to detect a neutrino

If ν are produced by β decay, they can be detected using the inverse reaction.



$\sigma(\nu p) \sim 10^{-43} \text{ cm}^2$ for $E_\nu = 3 \text{ MeV}$

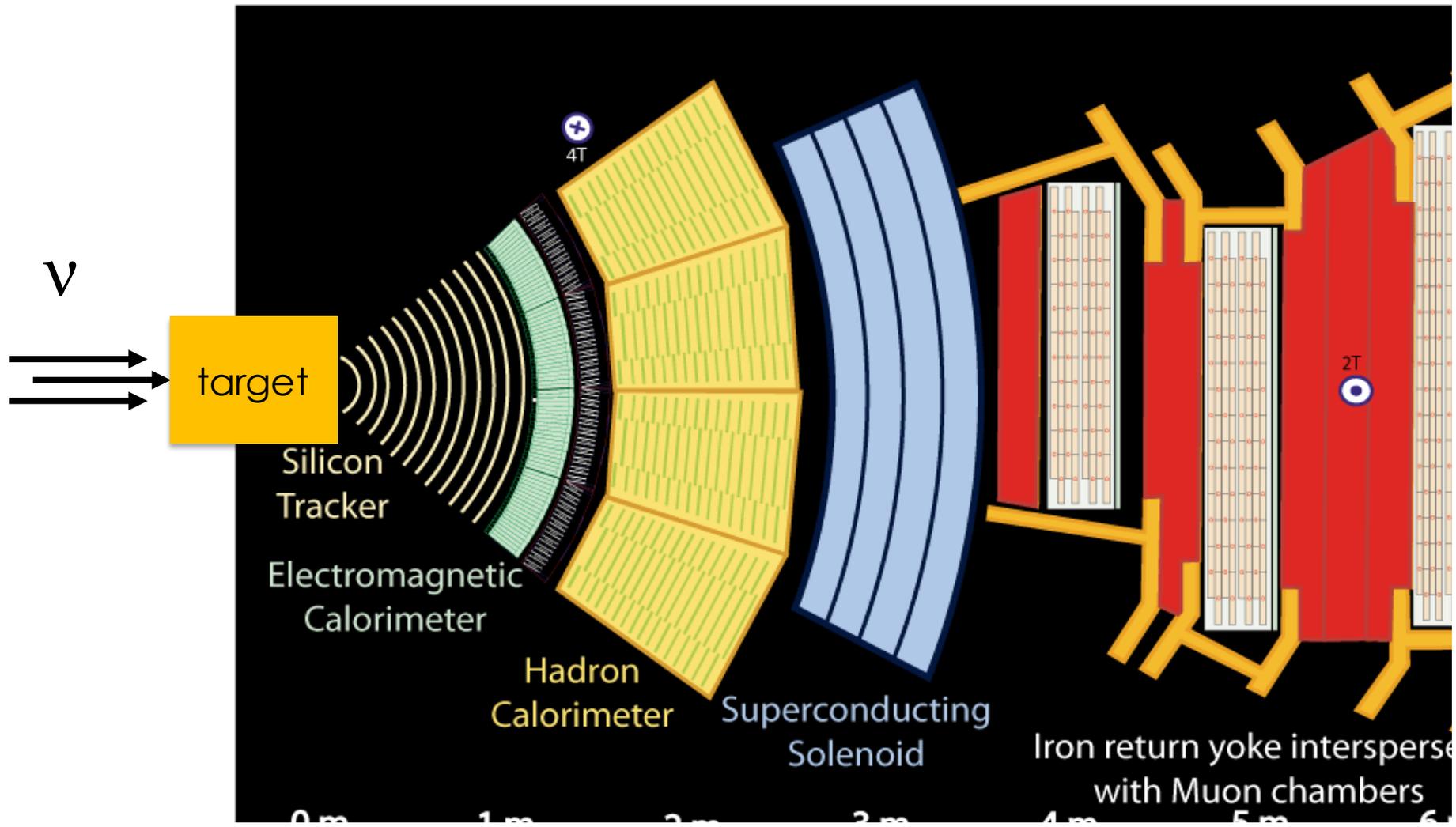
So one needs :

- Intense neutrino sources
- Large mass detectors

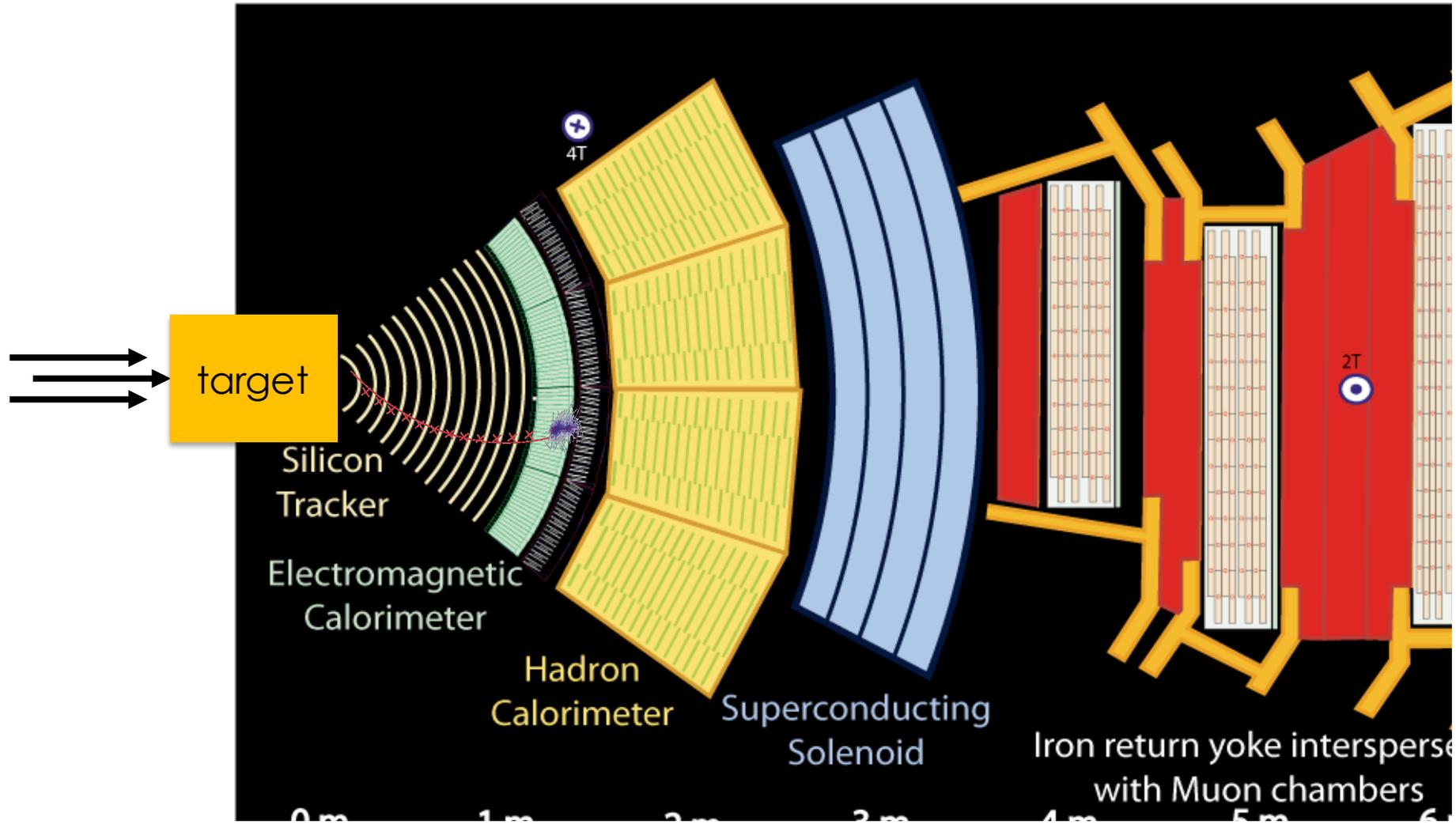
Sun
Cosmic rays interactions
Reactors
Accelerators



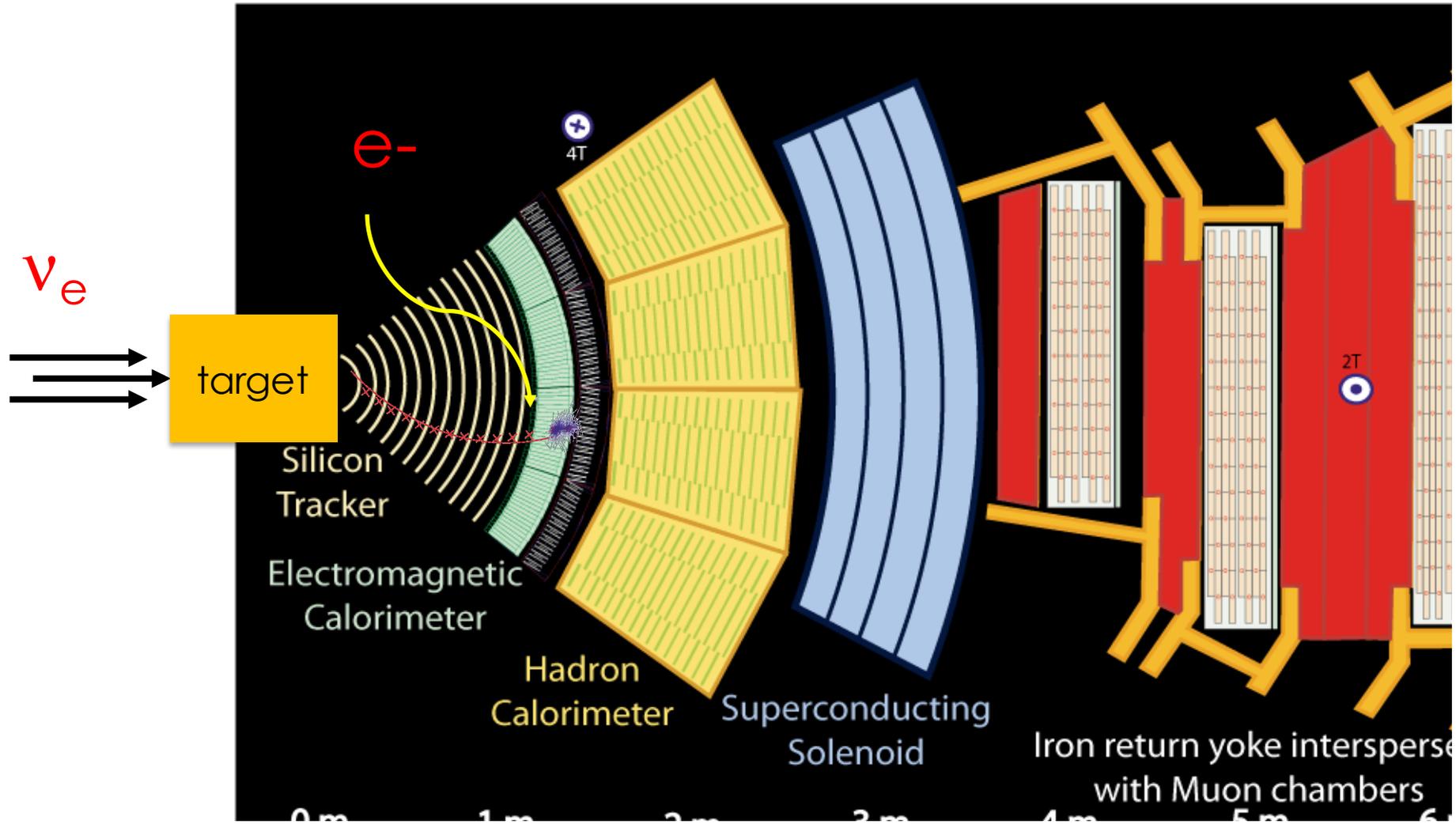
Three flavour of leptons : 3 flavour of neutrinos



Three flavour of leptons : 3 flavour of neutrinos

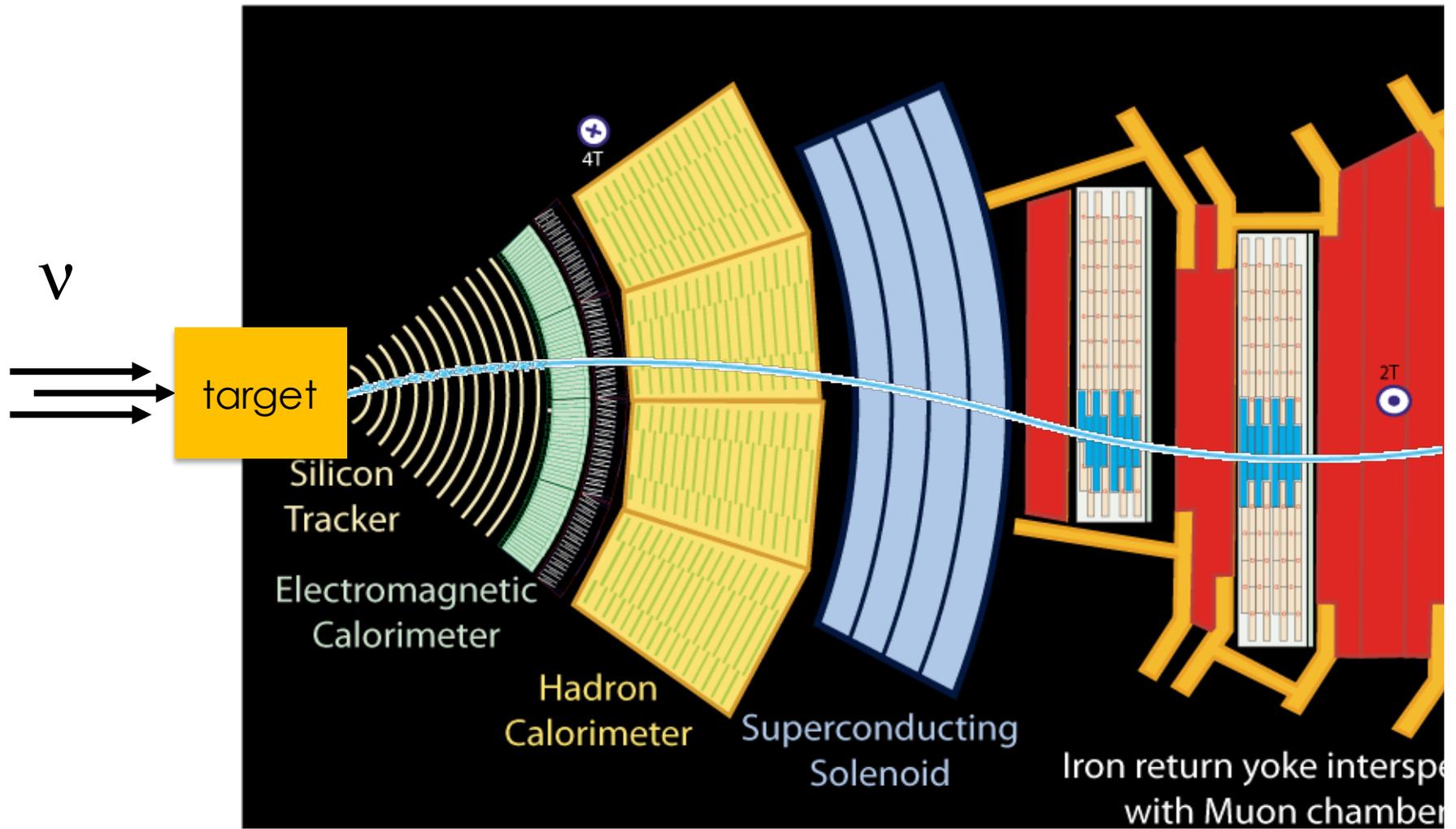


Three flavour of leptons : 3 flavour of neutrinos



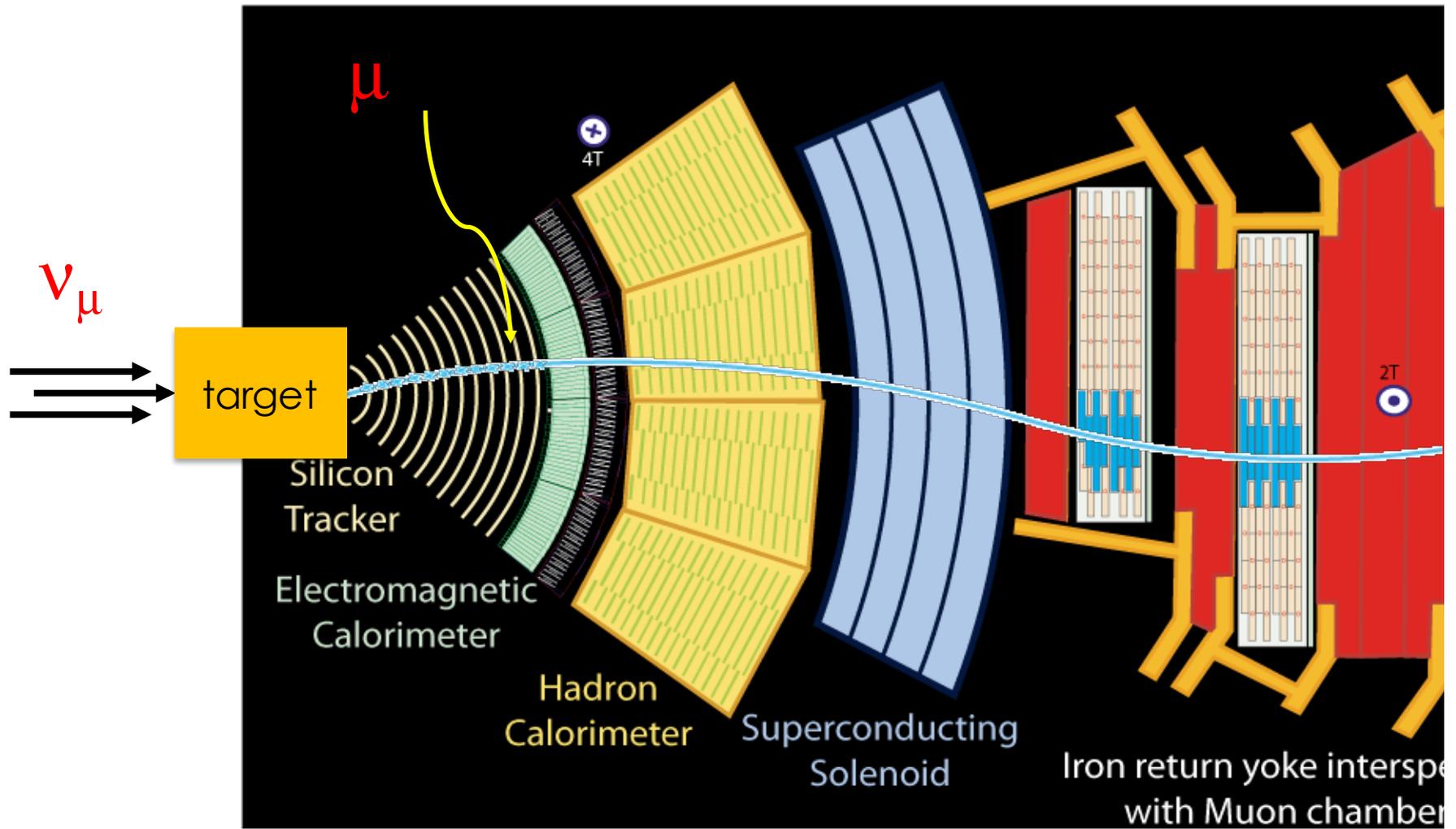
because of lepton number conservation

Three flavour of leptons : 3 flavour of neutrinos



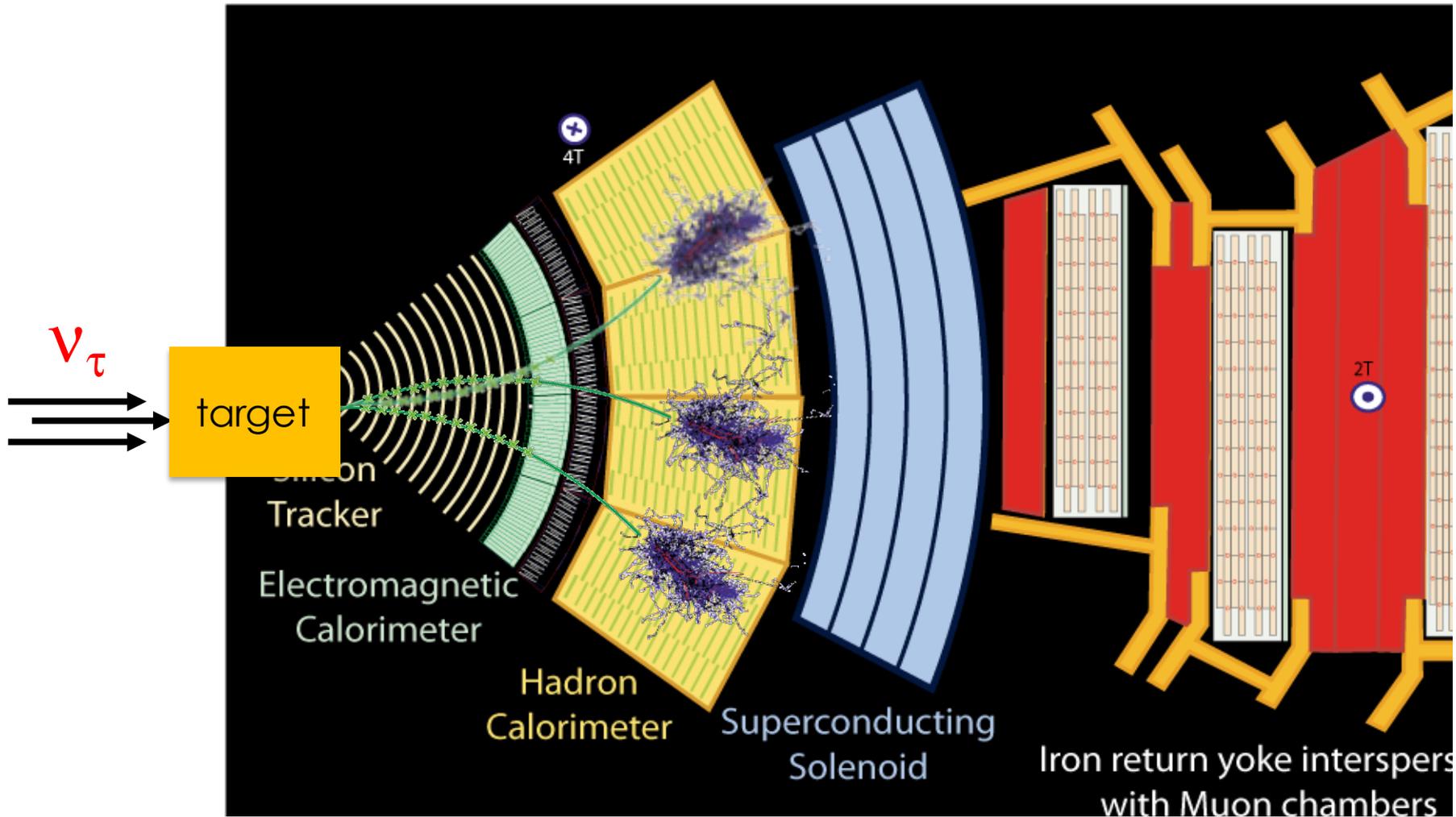
because of lepton number conservation

Three flavour of leptons : 3 flavour of neutrinos



because of lepton number conservation

Three flavour of leptons : 3 flavour of neutrinos

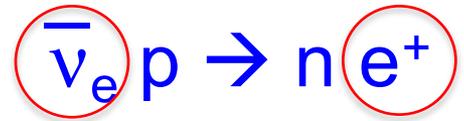


$$\tau^- \rightarrow \pi^- \pi^- \pi^+ \nu_\tau$$

because of lepton number conservation

Direct experimental evidence of $\bar{\nu}_e$

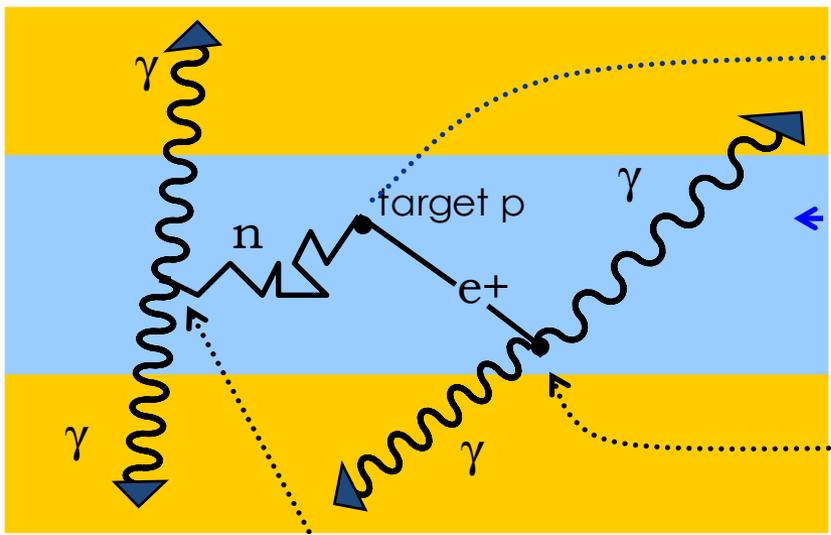
- Reminder: around 1930, Pauli and Fermi made the hypothesis of the ν_e
- In 1956, Reines-Cowan experiment : experimental evidence using a nuclear reactor : search for the $\bar{\nu}_e p \rightarrow n e^+$ reaction



Liquid scintillator

CdCl₂

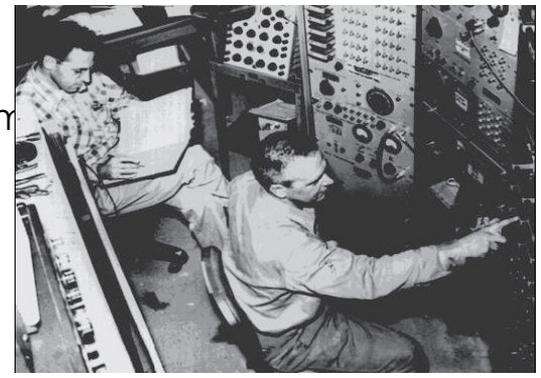
Liquid scintillator



$\bar{\nu}_e$ coming from the nuclear reactor

The e^+ stops and annihilates $e^+ e^- \rightarrow \gamma\gamma$

The neutron undergoes sequential collisions and after ~1 ms is captured by the Cadmium $\rightarrow \gamma$ emission

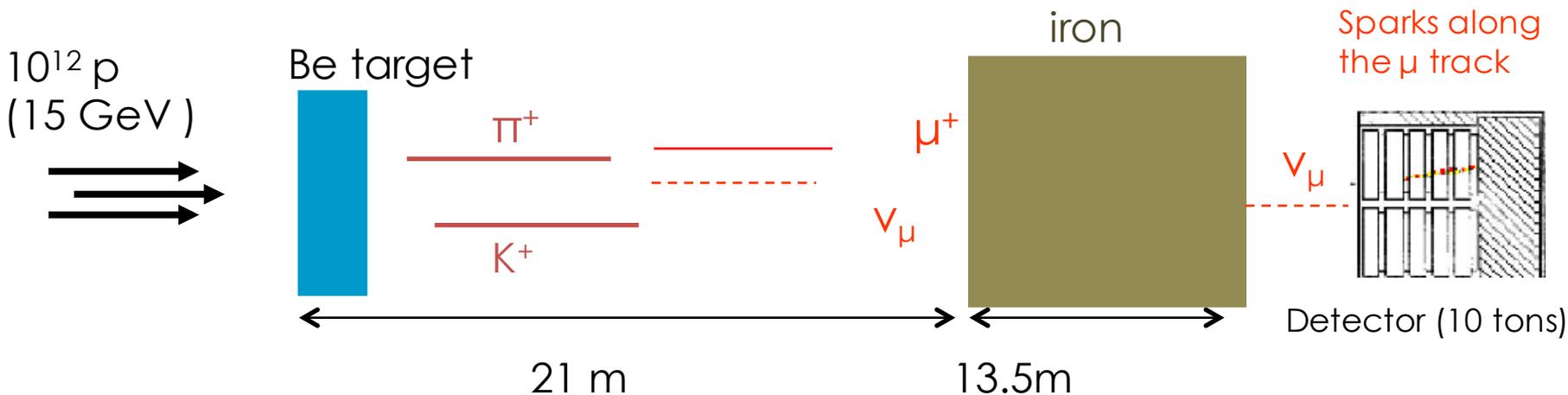


- Checks :
 - Target without cadmium
 - data taking with nuclear reactor OFF

$\rightarrow 3.0 \pm 0.2$ signal events/hour

Two other types de neutrinos

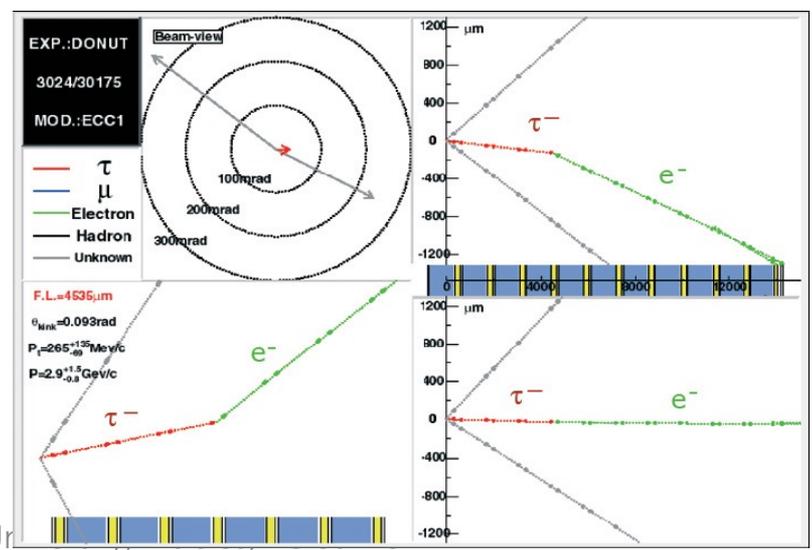
- In 1962 : Schwartz, Lederman et Steinberger experiment at BNL : $\nu_\mu \neq \nu_e$



- in 2000 the third neutrino ν_τ (DONUT) at Fermilab :



3 years of data taking : 4 unambiguous ν_τ signal events



BUT :

In a deep mine the Davis Chlorine experiment

Inverse beta decay $\text{Cl}^{37} + \nu \rightarrow \text{Ar}^{37} + e^-$  Searching for ν_e

Ar is chemically very different from Chlorine \rightarrow can be separated

It is radioactive and reverts to Cl^{37} emitting an Auger electron (lifetime 35 days)

Count a few atoms in a tank of few hundred tons !



The observed rate was about 3 times smaller than the predicted one ...

Which one is wrong ? The experiment or the Solar Model ?

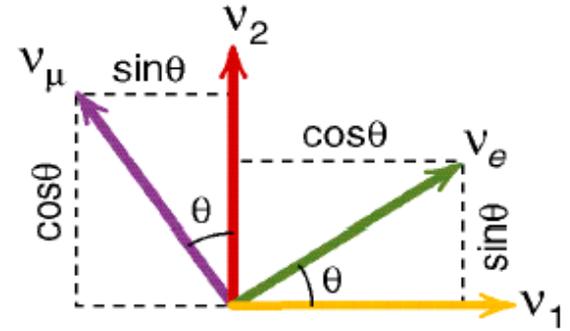
Neutrinos mixing

If neutrinos have masses : Mass eigenstates \neq weak (flavour) eigenstates

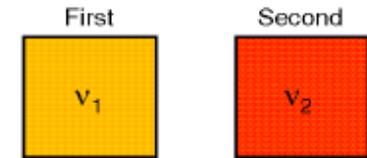
→ 3x3 matrix (CKM-style)

Assume for simplicity two families : the 2 basis are connected through a simple rotation :

$$\begin{pmatrix} \nu_e \\ \nu_\mu \end{pmatrix} = \begin{pmatrix} \cos \theta & \sin \theta \\ -\sin \theta & \cos \theta \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix} = U \begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix}$$



Mass states



Weak states



In the case of 3 families :

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

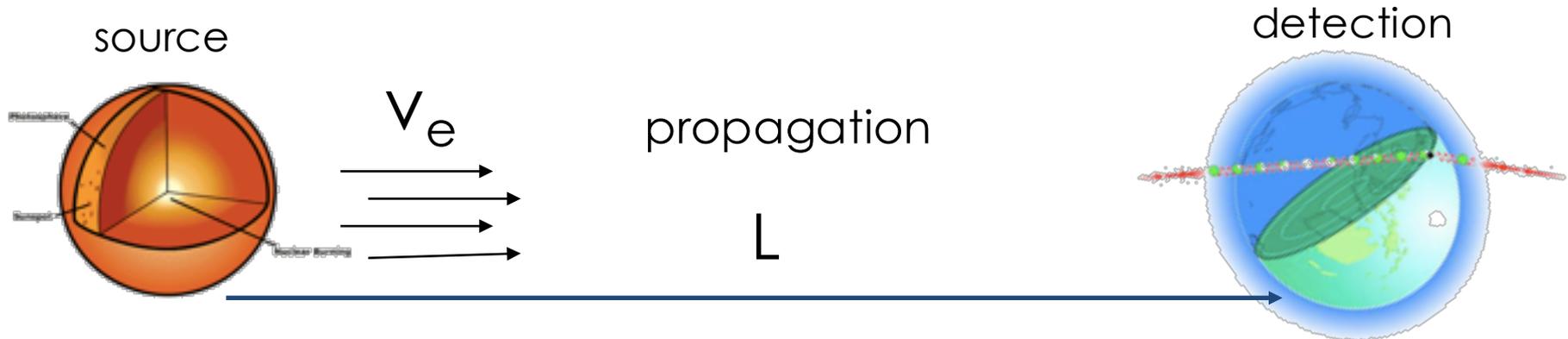
Weak interaction eigenstates (n_d)

PMNS matrix
(Pontecorvo, Maki, Nakagawa, Sakata)

Mass eigenstates (n_i)

$$\nu_\alpha = \sum_i^N U_{\alpha i} \nu_i$$

Neutrinos oscillations : the case for 2 families



The weak interaction produces neutrinos of a given flavor

$$|\nu(x=0)\rangle = |\nu_e\rangle = \cos\theta|\nu_1\rangle + \sin\theta|\nu_2\rangle$$

Evolution with time in the mass eigenstates basis follows Schrödinger equation (lab frame):

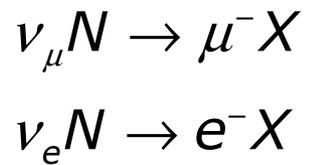
$$|\nu(L)\rangle = e^{-i(E_1t-p_1L)} \cos\theta|\nu_1\rangle + e^{-i(E_2t-p_2L)} \sin\theta|\nu_2\rangle$$

Ultra relativistic neutrinos of momentum p

$$p \gg M \Rightarrow E_i = \sqrt{p^2 + M_i^2} \approx p + \frac{M_i^2}{2p}$$

$$|\nu(L)\rangle = e^{-i\frac{M_1^2}{2p}L} \cos\theta|\nu_1\rangle + e^{-i\frac{M_2^2}{2p}L} \sin\theta|\nu_2\rangle$$

Detection again via weak interaction :



$$P(\nu_e \rightarrow \nu_\mu) = |\langle \nu_\mu | \nu(L) \rangle|^2 \approx \sin^2 2\theta \sin^2 \frac{\Delta m_{12}^2 L}{4E}$$

$$\Delta m_{12}^2 = M_1^2 - M_2^2$$

E is the average energy of the mass eigenstates

The **probability** for a ν_μ to transform into a ν_e at a distance L from the source is:

$$P(\nu_\mu \rightarrow \nu_e) = \sin^2(2\theta) \sin^2 \frac{(m_2^2 - m_1^2)L}{4E}$$

Mixing angle between flavor and mass states

Neutrino masses

$$P_{\alpha\beta} = \sin^2 2\theta \sin^2 \left(\frac{\Delta m^2 L}{4E_\nu} \right)$$

appearance

2 parameters
1 measurement

$$P_{\alpha\alpha} = 1 - P_{\alpha\beta} < 1$$

disappearance

The other flavour eigenstates (not present at $t=0$) appear during the propagation

The total flux is conserved.

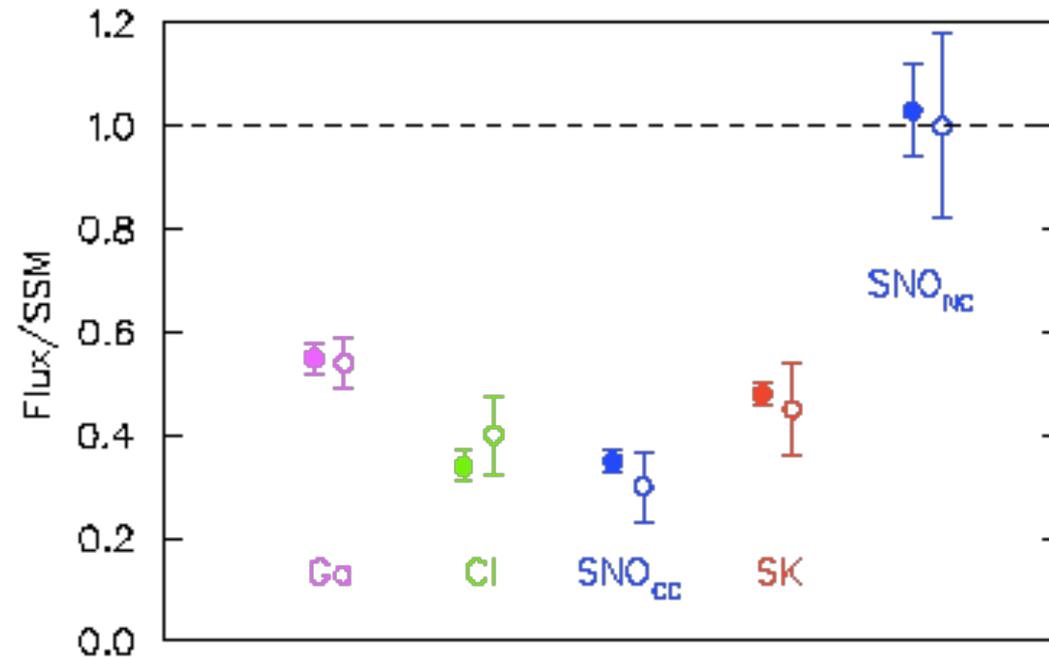
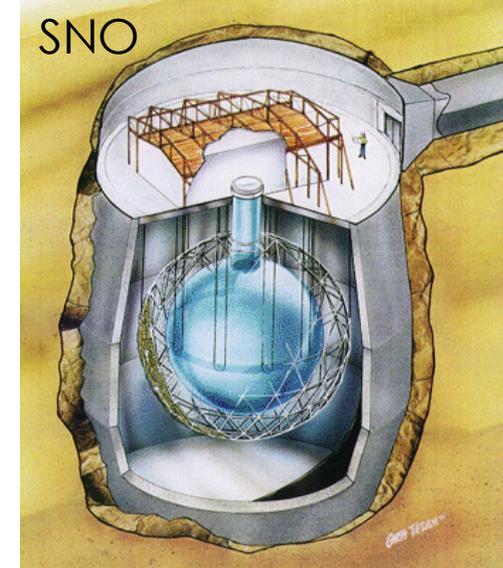
Neutrinos oscillations are only sensitive to the difference of the masses squared (NOT to the absolute value)

The mixing phenomenon can be tested using :

- Neutrinos from the sun
- Neutrinos from reactors
- Neutrinos from accelerators

If neutrinos oscillation is observed it implies that neutrinos have masses ...

Neutrinos from the sun !



the neutral current reaction (NC) measures the $(\nu_e + \nu_\mu + \nu_\tau)$ flux

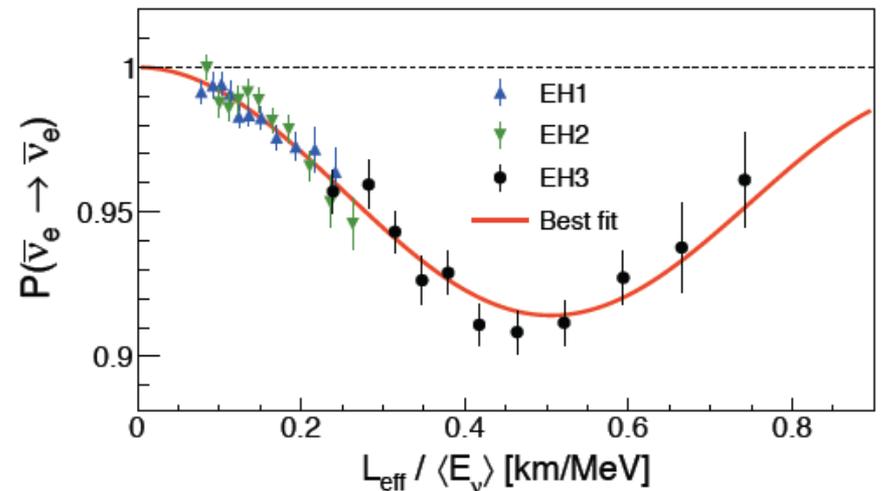
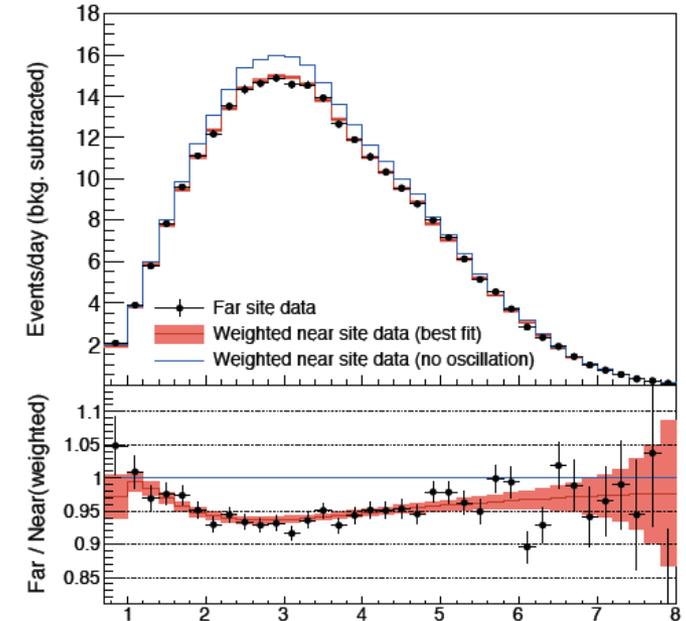
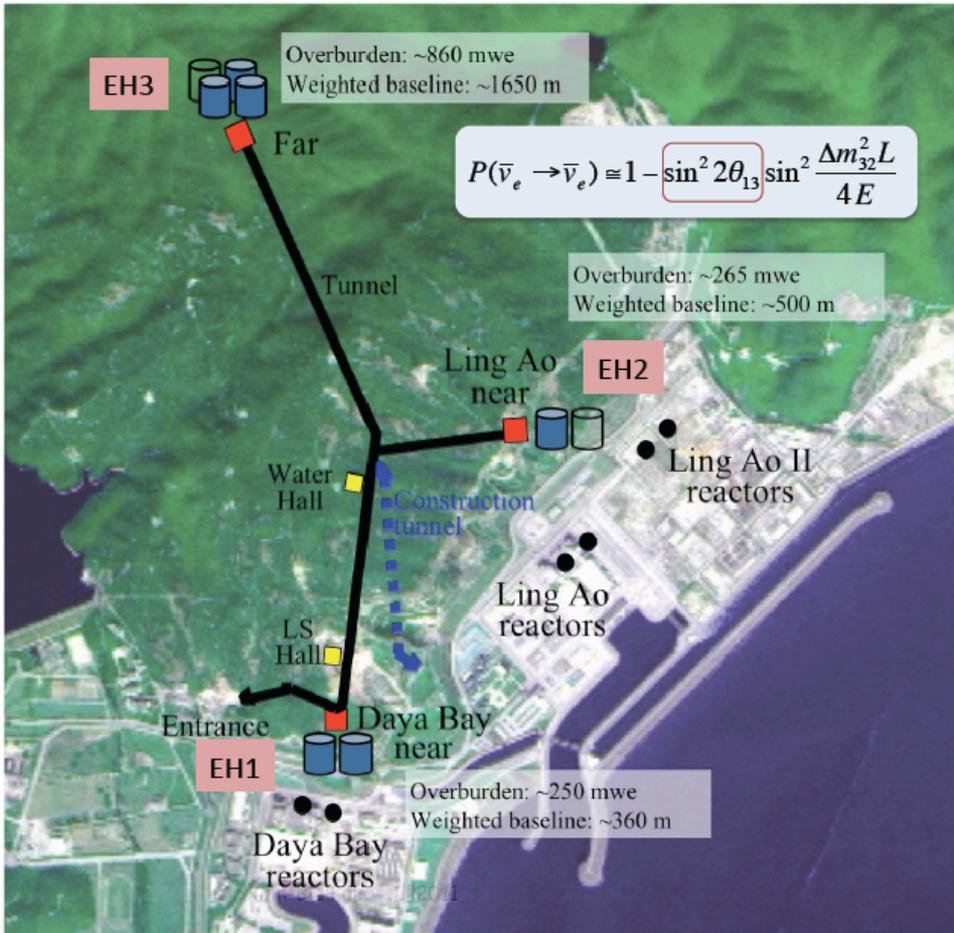
The charged current (CC) one the ν_e flux

The combination of all solar neutrino experiments (before SNO) implied that solar neutrinos were disappearing between production (in the sun core) and detection in the earth.

So the sun is shining the expected number of neutrinos but many of them are detected as ν_μ and/or ν_τ !

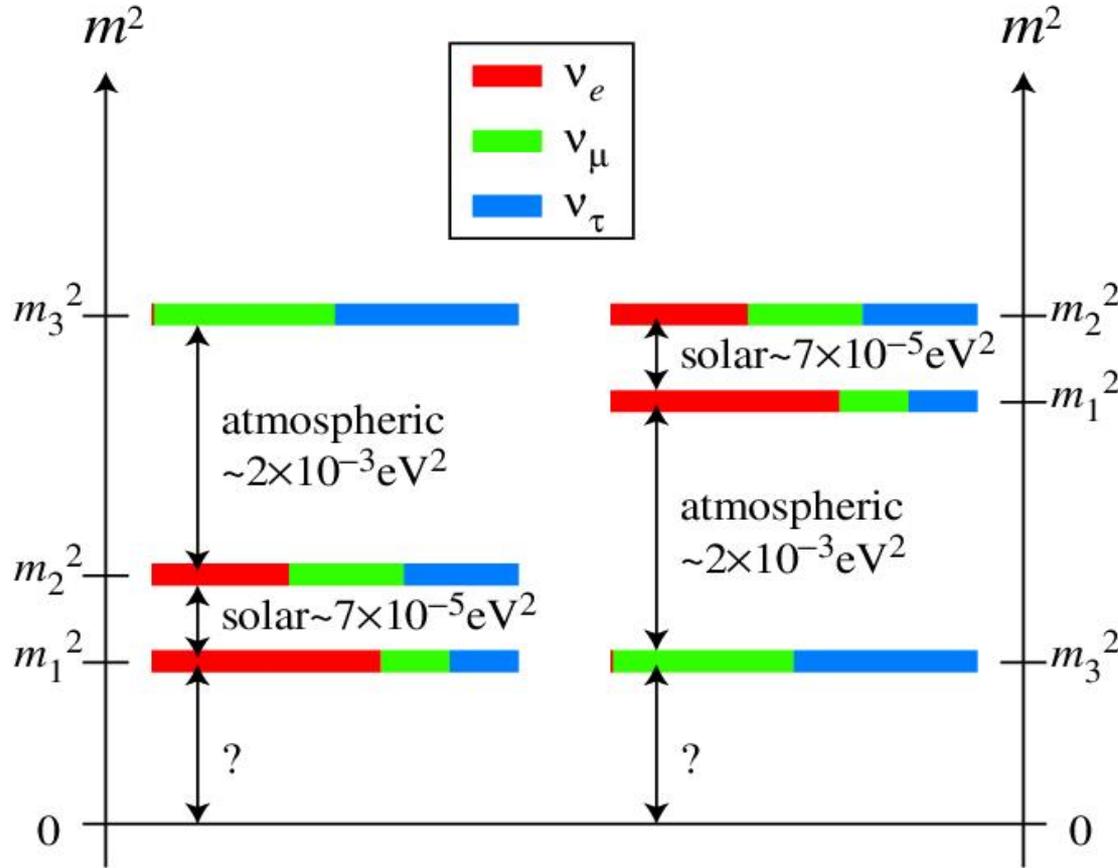
Neutrinos from the reactors : $\bar{\nu}_e$

PRL 115 111802 DayaBay



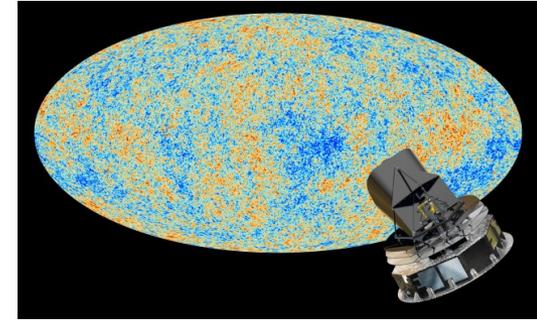
What is measured : $m_2^2 - m_1^2$ and $m_3^2 - m_2^2$

→ the neutrinos have masses but two scenarii are possible :



Which is the correct hierarchy ?

Cosmological constraints on $\Sigma m(\nu)$



- The neutrinos mass would modify the (delicate) balance between gravity and the Hubble expansion
- They would also have also effect on the structure formation ...
- Small modifications in the Cosmological Microwave Background which is the fingerprint of what happended at the very beginning of the Universe

$$\rightarrow \Sigma m(\nu) < \sim 0.2 \text{ eV}$$

Weak interaction in summary

- All quarks and leptons are sensitive to the weak interaction
- $M_W \sim M_Z \sim 100 \text{ GeV}$
 - short range
 - Extremely weak : ($\sim 10^{-8}$ smaller intensity than the strong interaction at a distance of 1 fm)
- The weak interaction
 - violates maximally C and P
 - does not conserve the flavour
 - Exhibits a tiny CP violation
- The weak and mass eigenstates of quarks are not the same, they are related via V_{CKM} which is a natural source of CP violation
- Neutrinos are only sensitive to the weak interaction
 - They have masses ! (but we do not know the values)
 - They may be an open window to physics beyond the Standard Model

3 families (3 angles (θ_{ij}) and one phase (δ)

$$V_{CKM} = \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13} \end{pmatrix} \quad V_{ub}$$

$$c_{ij} = \cos \theta_{ij}$$

$$s_{ij} = \sin \theta_{ij}$$

→ Parametrization in power of λ ($=\sin\theta_c$) = $s_{12} = |V_{us}| \sim 0.22$

$$\begin{pmatrix} 1 - \lambda^2/2 & \lambda & A\lambda^3(\rho - i\eta) \\ -\lambda & 1 - \lambda^2/2 & A\lambda^2 \\ A\lambda^3(1 - \rho - i\eta) & -A\lambda^2 & 1 \end{pmatrix} + \mathcal{O}(\lambda^4)$$

$$\lambda = \sin \theta_c \sim 0.22$$

$$A \sim 0.80$$

$$\rho \sim 0.20$$

$$\eta \sim 0.35$$

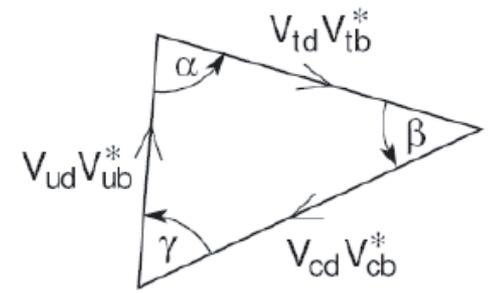
“the” unitarity triangle :

$$V_{ub}^* V_{ud} + V_{cb}^* V_{cd} + V_{tb}^* V_{td} = 0$$

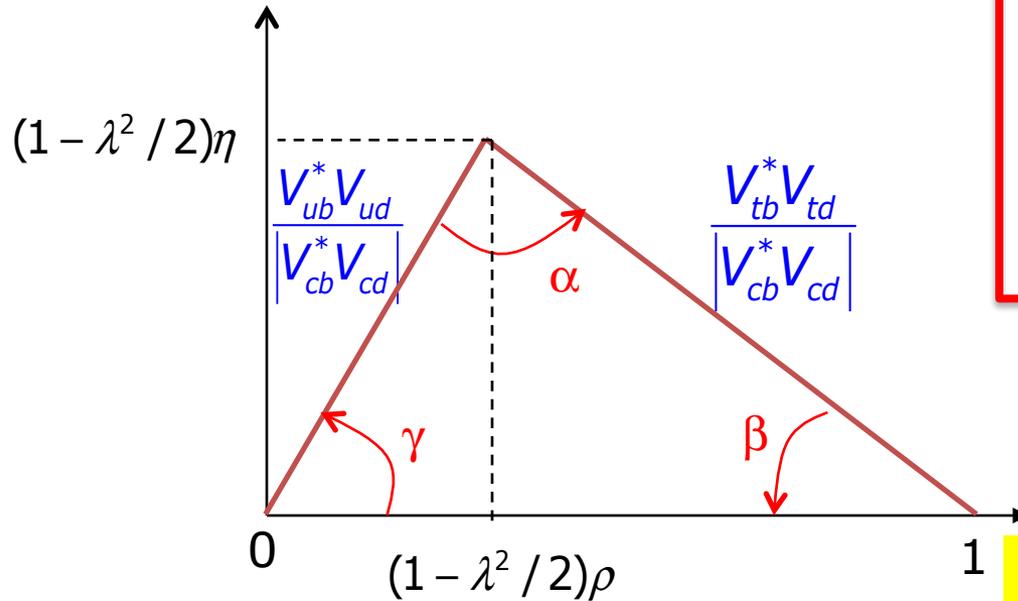
$$V_{td} V_{tb}^* = A\lambda^3(1 - \rho - i\eta) + A\lambda^5(\rho + i\eta)$$

$$V_{ud} V_{ub}^* = A\lambda^3(\rho + i\eta) \times \left(1 - \frac{\lambda^2}{2}\right) \quad \text{at order } \lambda^5$$

$$V_{cd} V_{cb}^* = -A\lambda^3$$



Basis of the triangle aligned on the real axis, normalized to 1



$$\beta = \arg\left(\frac{V_{td} V_{tb}^*}{|V_{cd} V_{cb}^*|}\right) = \text{atan}\left(\frac{(1 - \lambda^2/2)\eta}{1 - (1 - \lambda^2/2)\rho}\right)$$

$$\gamma = \arg\left(\frac{V_{ud} V_{ub}^*}{|V_{cd} V_{cb}^*|}\right) = \text{atan}\left(\frac{\eta}{\rho}\right)$$

$$\alpha + \beta + \gamma = \pi$$

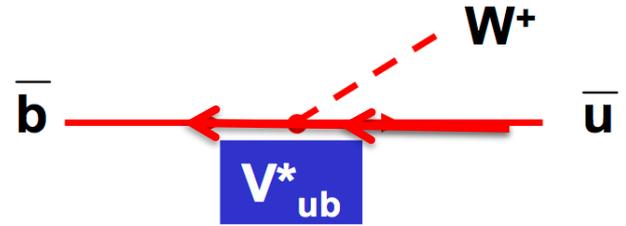
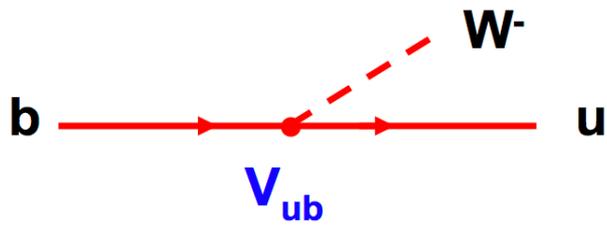
2 sides ; 3 angles

⇒ aim : to overconstrain this **unitarity triangle**

precision test of the Standard Model



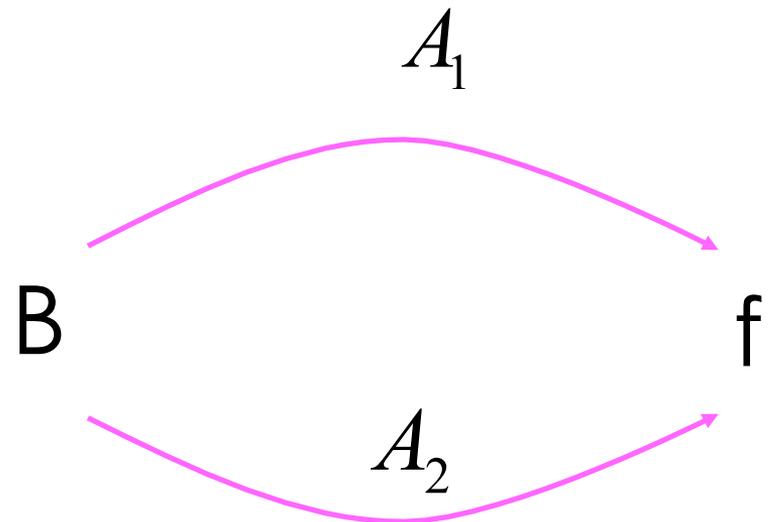
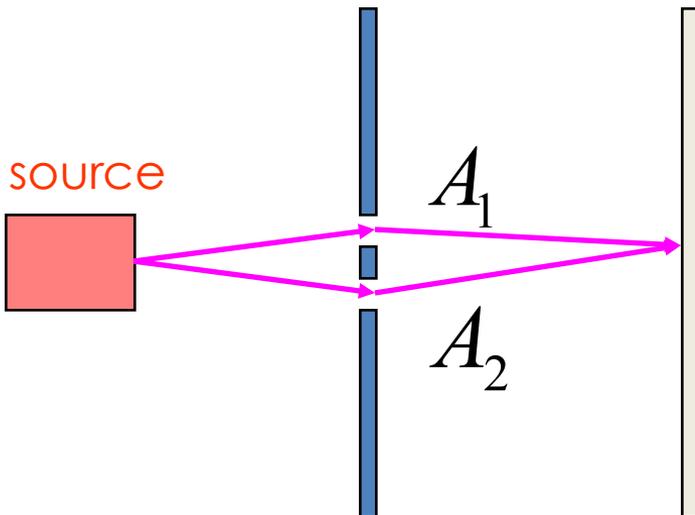
CP violation



$V_{ub}^* \neq V_{ub} \rightarrow$ CP violation

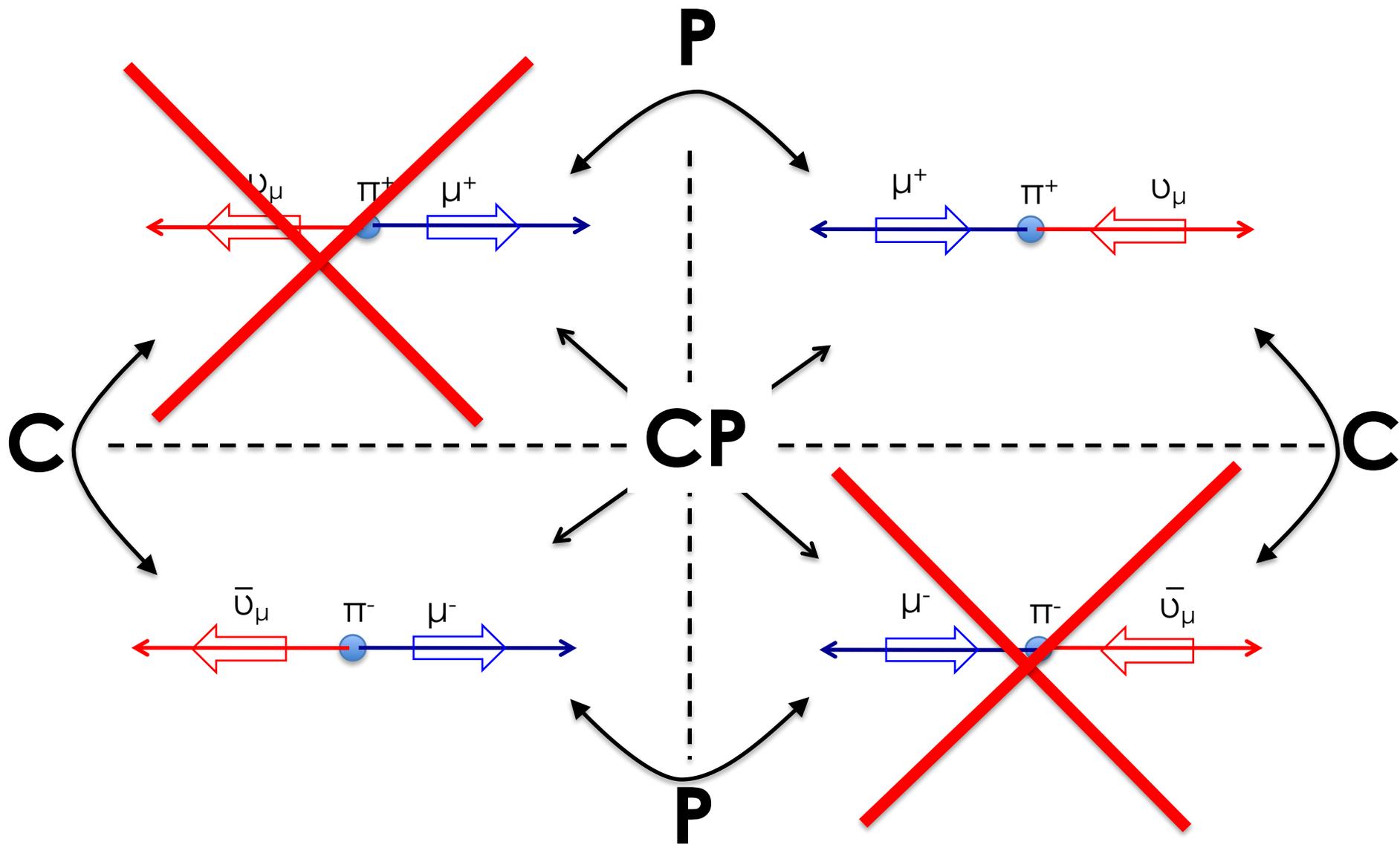
CP

If you just have one amplitude : no sensitivity on phase ($|V_{ij}|^2 = |V_{ij}^*|^2$)

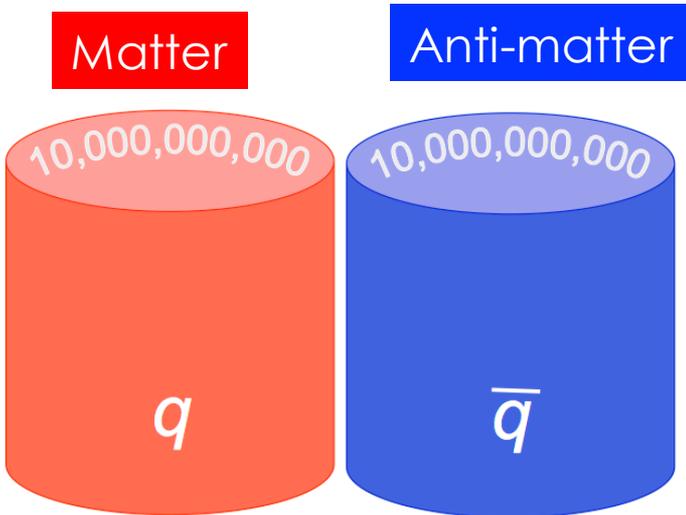




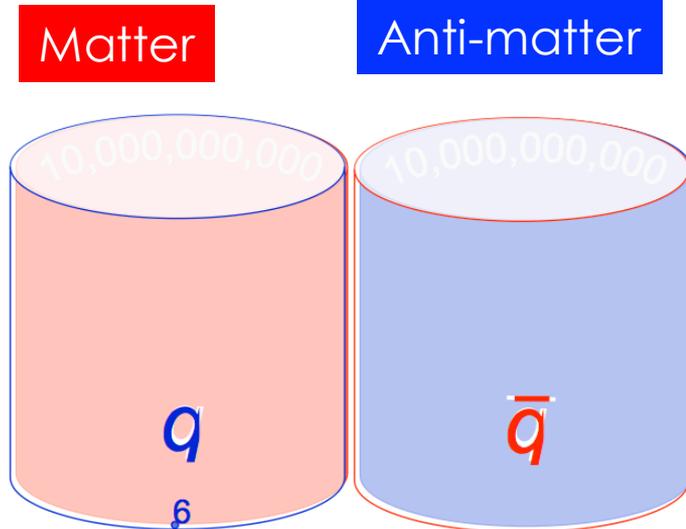
Let's come back to the unitarity triangle



Anti-matter in the Universe and Big Bang



Primordial Universe



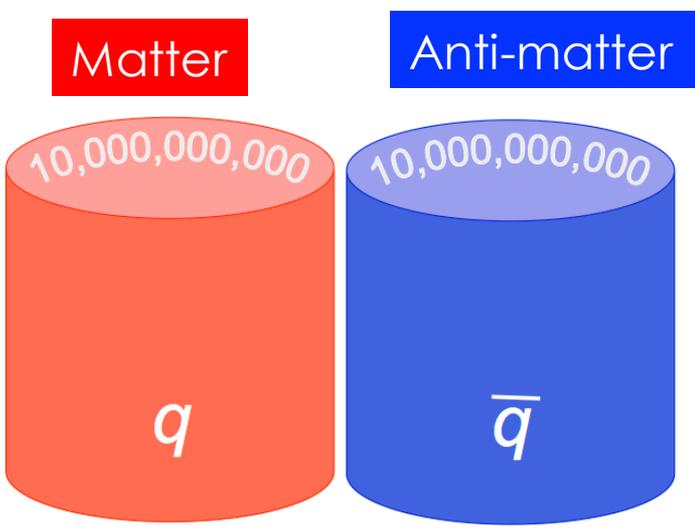
Today

$$\frac{n(\text{baryon}) - n(\text{antibaryon})}{n_\gamma} \sim 6 \cdot 10^{-10}$$

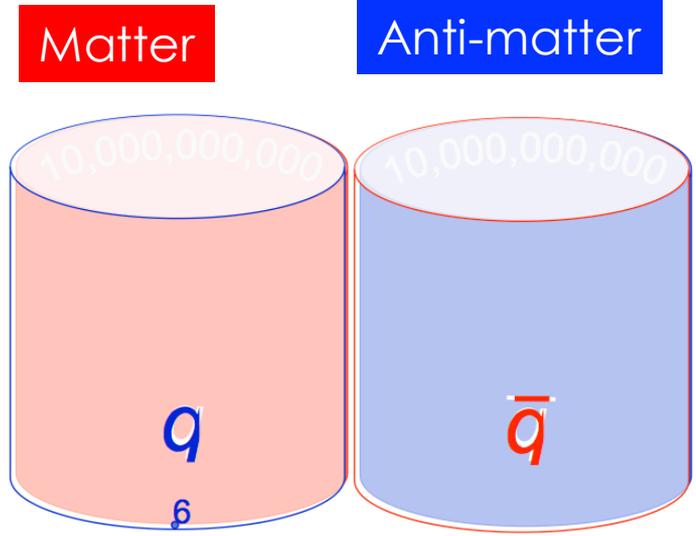
The 3 Sakharov conditions(1967)

1. Baryonic number violation: $X \rightarrow p e^-$
2. C and CP symmetries violation: $\Gamma(X \rightarrow p e^-) \neq \Gamma(\bar{X} \rightarrow \bar{p} e^+)$
3. To be out of equilibrium: $\Gamma(X \rightarrow p e^-) \neq \Gamma(p e^- \rightarrow X)$

Anti-matter in the Universe and Big Bang



Primordial Universe



Today

$$\frac{n(\text{baryon}) - n(\text{antibaryon})}{n_\gamma} \sim 6 \cdot 10^{-10}$$

The 3 Sakharov conditions(1967)

1. Baryonic r
2. C and CP
3. To be out of equilibrium.

But the CP violation phase of the SM is orders of magnitude too small